



Science and
Technology
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UK Astronomy
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The South African Prototype SRC

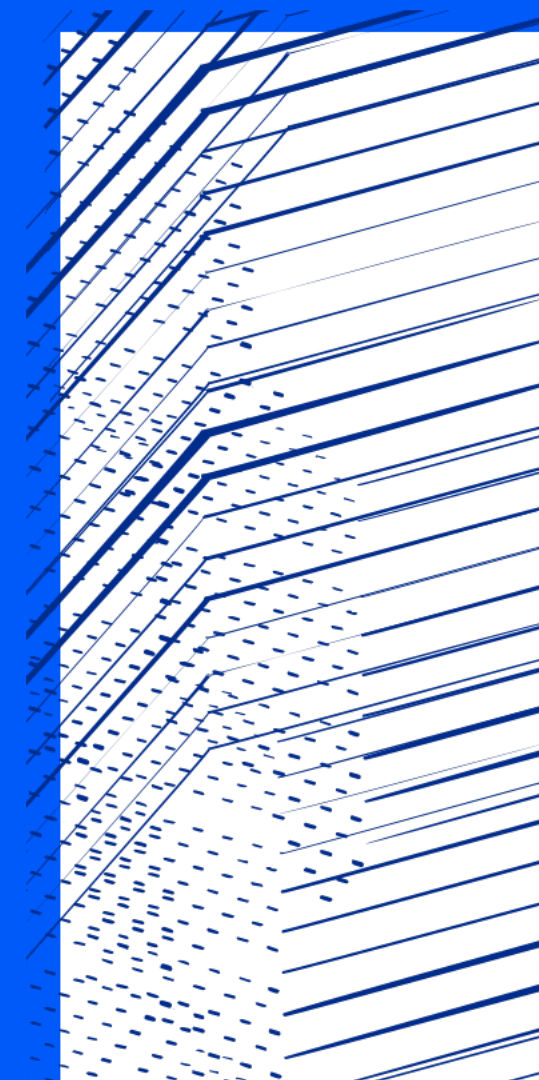
IDIA and MeerKAT

Dr. Bradley Frank

Head of Electronic and Software Engineering — UKATC

Formerly SARA0's Associate Director for Astronomy Operations at IDIA

Acknowledging Russ Taylor (Director) and Rob Simmonds (AD Tech)



MeerKAT

- 2006: ~20-dish Karoo Array Telescope



#1 / 2006

This issue features:

Minister visits proposed astronomy reserve

Four competing countries to present their SKA

How the KAT is taking shape

KAT funding and astronomy reserve on track

MeerKAT

- 2006: ~20-dish Karoo Array Telescope
- 2008: 80x12m *meerKAT*



#2 / 2007

This issue features:

MeerKAT prototype antenna installed

Feeding the meerKAT prototype

European award for SA astronomy project

SKA champions in South Africa



The meerKAT prototype with all major mechanical components installed. The prototype is at the Hartebeesthoek Radio Astronomy Observatory in South Africa.



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Technology Centre



SKA Africa eNews

October 2010

Inside this issue

- 1 Innovative new design for SA's MeerKAT
- 2 New measurements of radio frequency interference at the proposed South African and Australian SKA sites
- 3 Africa unites behind continent's SKA bid
- 4 IAU Astronomy Development Office will be in South Africa
- 4 Developing skills for the future
- 5 Global partnerships to meet SKA demands

Innovative new design for SA's MeerKAT

Following an extensive engineering design process, the baseline design concept for the South African MeerKAT precursor telescope has been decided. This design process consisted of an in-depth design study that investigated implementation options and tradeoffs for all key subsystems, and culminated in a Concept Design Review (CoDR) undertaken by an independent panel of international experts. The recommendations of the CoDR panel have informed the baseline concept, and the most visible design decision is that the MeerKAT will consist of 64 Gregorian offset dishes, each with an effective diameter of 13.5 metres.

"This is the most innovative option of the design solutions that we considered, but it will allow the MeerKAT to operate at a sensitivity of over 220 m²/K" explains Anita Loots, Associate Director of the SKA South Africa Project.

With all seven dishes of the MeerKAT precursor array (known as KAT-7) now in place, the construction of MeerKAT itself is the next big step for the SKA Africa team. "We will start by building a qualification (prototype) dish of the new design, on site in the Karoo," Loots adds. This first dish will be located near the KAT-7 array, which will allow extensive testing of the performance of the new design against the existing array.

(CoDR) undertaken by an independent panel of international experts. The recommendations of the CoDR panel have informed the baseline concept, and the most visible design decision is that the MeerKAT will consist of 64 Gregorian offset dishes, each with an effective diameter of 13.5 metres.

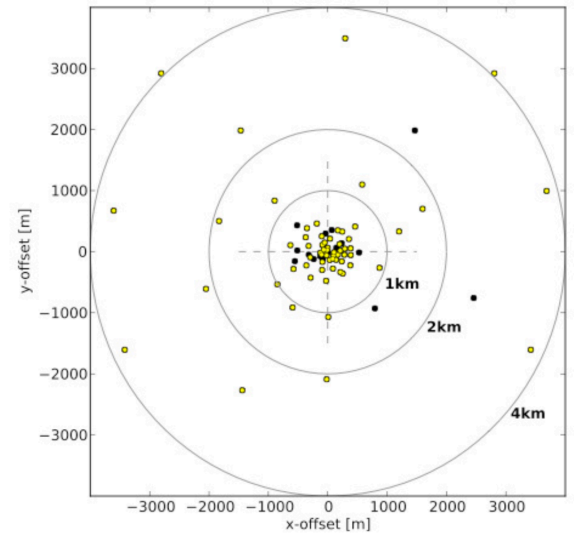


Figure 27: 64-dish Ver. 3.6 (yellow only) full configuration. Yellow and black together form the 80-dish Ver. 3.6

Commercial in Confidence

M0000-0000V1-03 DD
Revision: 1

An iteration of structural design was performed which resulted in the layout presented in Figure 3.

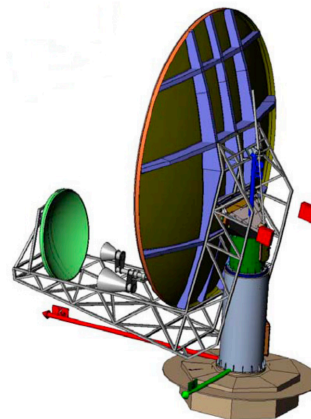


Figure 3: Option 3 – structural layout

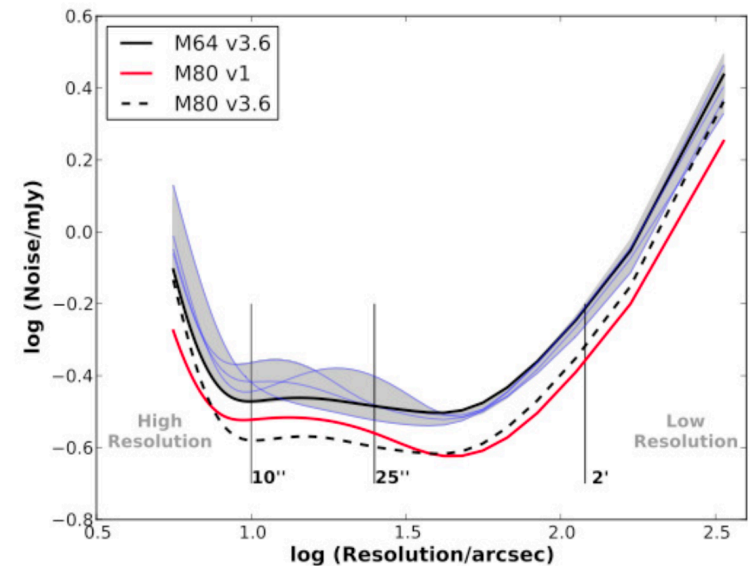


Figure 26: Sensitivity performance versus resolution for array configurations



UK Astronomy Technology Centre

From the MeerKAT Configuration Design Review, 2010

MeerKAT Large Science Projects



An open invitation to the Astronomical Community
to propose Key Project Science with the South African
Square Kilometre Array Precursor

MeerKAT

R.S. Booth

Hartebeesthoek Radio Astronomy Observatory, P.O.Box 443, Krugersdorp 1740, South Africa
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J.L. Jonas

Rhodes University, Dept. Physics & Electronics, PO Box 94, Grahamstown 6410, South Africa
email: j.jonas@ru.ac.za

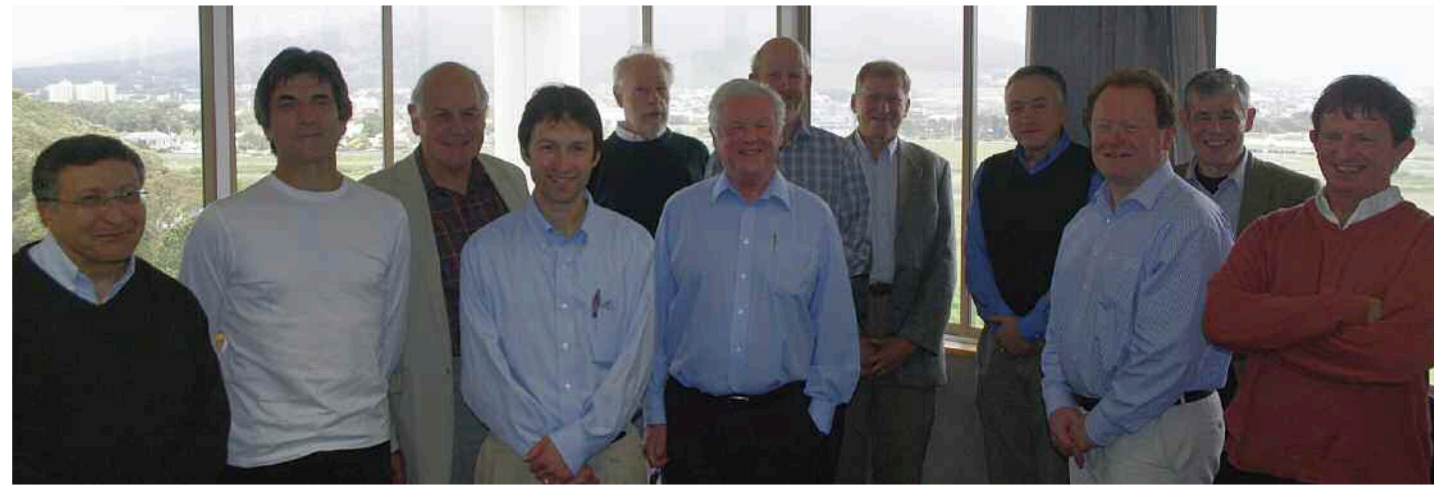
B. Fanaroff

SKA South Africa Project Office, 17 Baker St, Rosebank, Johannesburg, South Africa
email: bfanaroff@fanaroff.co.za

Proposal Submission deadline: March 15, 2010

1 Introduction

As possible hosts of the Square Kilometre Array (SKA), South Africa and Australia



Large Survey Projects

A call for large MeerKAT observing proposals by SRAO's predecessor SKA South Africa in 2009 resulted in 10 Large Survey Projects (LSPs, defined to require more than 1000 hours of telescope time over 5 years) being approved in 2010. In 2016, with MeerKAT in an advanced state of construction, SKA South Africa requested revised project plans from eight LSPs:

- [LADUMA](#): Looking at the Distant Universe with the MeerKAT Array
- [MALS](#): The MeerKAT Absorption Line Survey
- [MeerKAT Fornax Survey](#)
- [MeerTime](#): The MeerKAT Key Science Project on Pulsar Timing
- [MHONGOOSE](#): MeerKAT HI Observations of Nearby Galactic Objects: Observing Southern Emitters
- [MIGHTEE](#): The MeerKAT International GHz Tiered Extragalactic Exploration Survey
- [ThunderKAT](#): The Hunt for Dynamic and Explosive Radio Transients with MeerKAT
- [TRAPUM](#): Transients and Pulsars with MeerKAT <https://skafrica.atlassian.net>

MIGHTEE

- Wide-band, wide-field survey at L-band.
- Four footprints, overlapping with large multiwavelength surveys.

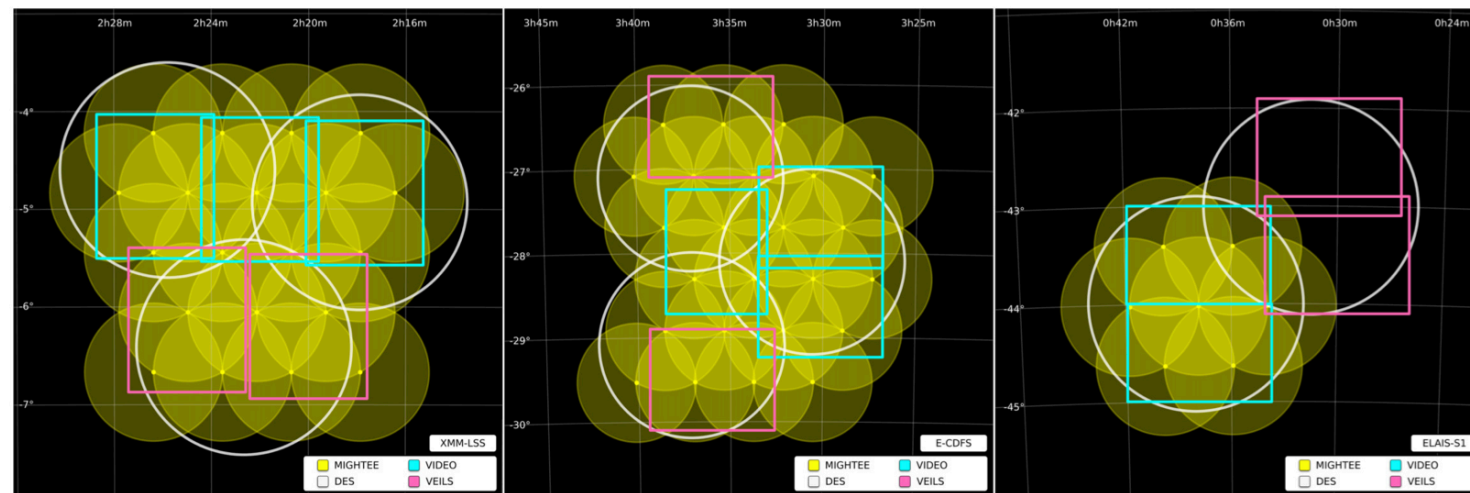


Figure 4: Current plausible pointing strategies for (left-to-right) XMM-LSS (20 pointings, 6.7 deg^2), E-CDFS (24 pointings, 8.3 deg^2) and ELAIS-S1 (7 pointings, 1.6 deg^2). Not shown here is the fourth COSMOS field, which will consist of a single deep pointing. In practice the grid of E-CDFS pointings will be snapped to the LADUMA pointing centre, requiring only 23 additional pointings from MIGHTEE.

MIGHTEE-HI


- Science Goals — Maddox et al. (2020)
 - HI Mass & Velocity Function
 - Kinematics — Spatially resolved HI.
 - Tully-Fisher Relation.
 - HI in Low Mass Galaxies.
 - HI as a function of Environment.
 - HI below the detection limit.
 - Galaxies and the surrounding medium.

Free Access

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Published online	02 February 2021

A&A 646, A35 (2021)

MIGHTEE-HI: The HI emission project of the MeerKAT MIGHTEE survey*

 N. Maddox¹, B. S. Frank^{2,3,4}, A. A. Ponomareva⁵, M. J. Jarvis^{5,6}, E. A. K. Adams^{7,8}, R. Davé^{9,6,10}, T. A. Oosterloo^{7,8}, M. G. Santos^{6,2}, S. L. Blyth⁴, M. Glowacki¹¹, R. C. Kraan-Korteweg⁴, W. Mulaudzi⁴, B. Namumba¹², I. Prandoni¹³, S. H. A. Rajohnson⁴, K. Spekkens¹⁴, N. J. Adams⁵, R. A. A. Bowler⁵, J. D. Collier^{3,15}, I. Heywood^{5,12,2}, S. Sekhar^{3,16} and A. R. Taylor^{3,11}



Received: 12 October 2020 | Accepted: 10 November 2020

Abstract

We present the HI emission project within the MIGHTEE survey, currently being carried out with the newly commissioned MeerKAT radio telescope. This is one of the first deep, blind, medium-wide interferometric surveys for neutral hydrogen (HI) ever undertaken, extending our knowledge of HI emission to $z = 0.6$. The science goals of this medium-deep, medium-wide survey are extensive, including the evolution of the neutral gas content of galaxies over the past 5 billion years. Simulations predict nearly 3000 galaxies over $0 < z < 0.4$ will be detected directly in HI, with statistical detections extending to $z = 0.6$. The survey allows us to explore HI as a

MeerKAT Science: On the Pathway to the SKA

MeerKAT2016

25-27 May, 2016
Stellenbosch, South Africa
published February 01, 2018

[Entries on ADS](#)

MeerKAT Science: On the Pathway to the SKA

MeerKAT is a next generation radio telescope under construction on the African SKA central site in the Karoo plateau of South Africa. When completed in 2018 MeerKAT will be a 64-element array of 13.5-m parabolic antennas distributed over an area with a diameter of 8 km. With a combination of wide bandwidth and field of view, with the large number of antennas and total collecting area, MeerKAT will be one of the world's most powerful imaging telescopes operating at GHz frequencies.

MeerKAT is a science and technology precursor of the SKA mid-frequency dish array, and following several years of operation as a South African telescope will be incorporated into the SKA phase-one facility. The MeerKAT science program will consist of a combination of key science, legacy-style, large survey projects, and smaller projects based on proposals for open time. This workshop, which took place in Stellenbosch in the Western Cape, was held to discuss and plan the broad range of scientific investigations that will be undertaken during the pre-SKA phase of MeerKAT. Topics covered included:

- technical development and roll out of the MeerKAT science capabilities,
- details of the large survey projects presented by the project teams,
 - science program concepts for open time,
- commensal programs such as the Search for Extraterrestrial Intelligence,
- and the impact of MeerKAT on global Very Long Baseline Interferometry.

Editorial Board

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Fernando Camilo
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Kavilan Moodley

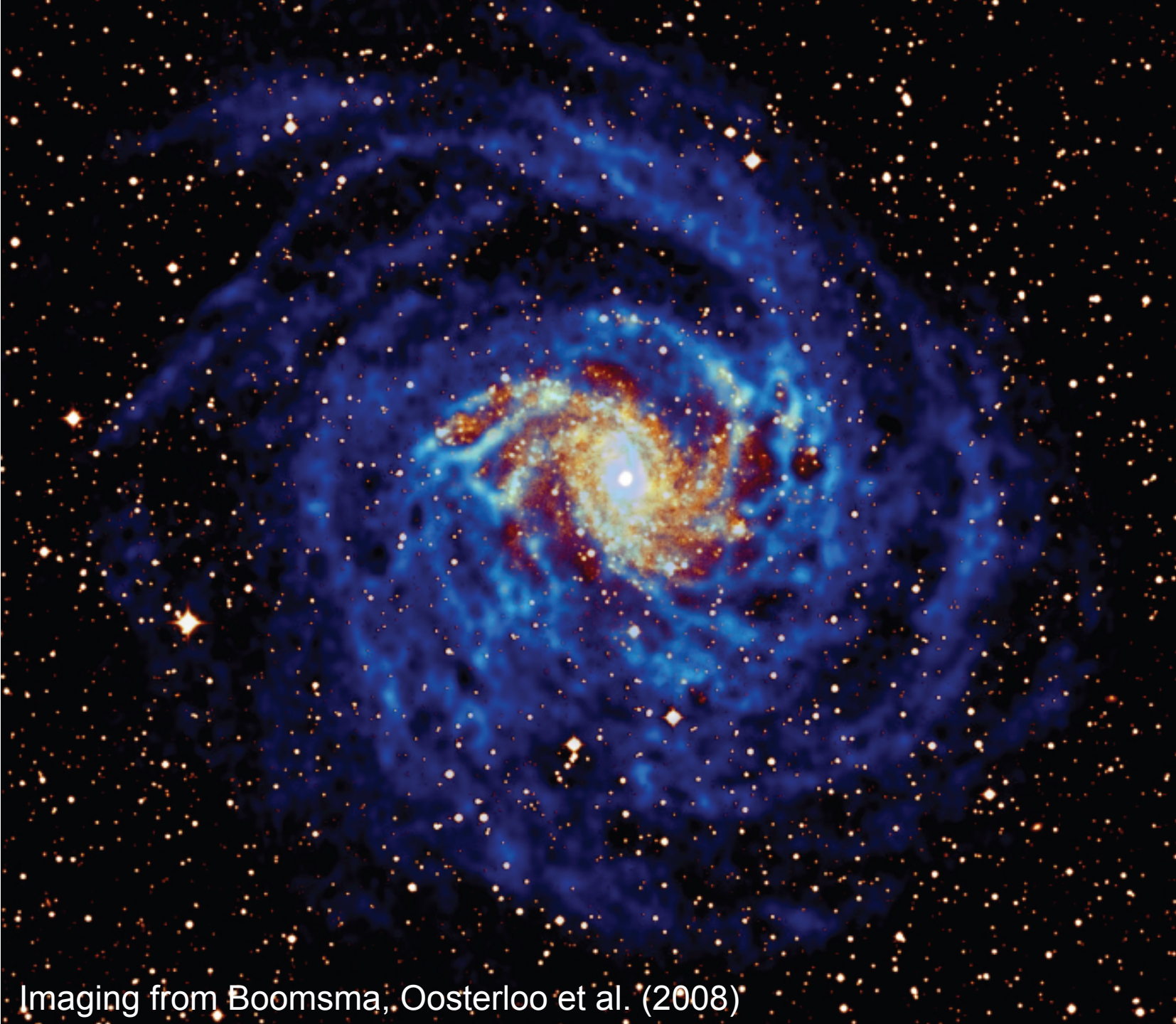
Report of the MeerKAT Large Project Review Panel – 2017

I. Background and Motivation for a Review

In 2009 the Square Kilometre Array (SKA) South Africa (SA) office released an open invitation to the international astronomical community, calling for key science projects to be carried out on the SKA precursor MeerKAT (Booth et al. astro-ph/0910.2935). A committee was formed in 2010 chaired by J. Lazio (JPL) to review these proposals and recommend a large science program for MeerKAT. The committee chose ten Large Survey Projects (LSPs), defined as requiring more than 1000 hours of telescope time over five years, and expected approximately 70% of the observing time to be allocated to LSPs during that span.

Table 1. Rank-ordered list of the recommended MeerKAT large science program.

Large Survey Project (LSP) Components	Requested Time (hrs)	Readers' Science Score	Panel Ranking
MeerTime (binary)	1440	3.87	A
MHONGOOSE	1650	3.55	A
MeerTime (MSPs)	2160	3.58	A
LADUMA	3424	3.84	A
Fornax	900	3.41	A
TRAPUM (Fermi sources)	338	3.60	A
MeerTime (1000 PTA)	720	3.78	A/B
ThunderKAT (CVs)	250	3.42	B
MIGHTEE (L band)	979	3.23	B
ThunderKAT (GRBs)	330	3.42	B
MeerTime (GCs)	1080	3.38	B
MALS (UHF band)	2320	N/A	B
TRAPUM (nearby galaxies)	226	3.28	B
TRAPUM (GCs)	320	3.22	B
TRAPUM (SNR, PWN, TeV)	92	N/A	B
ThunderKAT (SNe Ia)	200	3.08	B
MIGHTEE (S band)	948	2.77	B/C
MeerTime (magnetars)	100	3.00	C
TRAPUM (Fly's Eye)	720	3.33	C
ThunderKAT (XRBs)	500	3.00	C
MALS (L band)	1073	N/A	C
MeerTime (young PSRs)	400	2.83	C



The benchmark

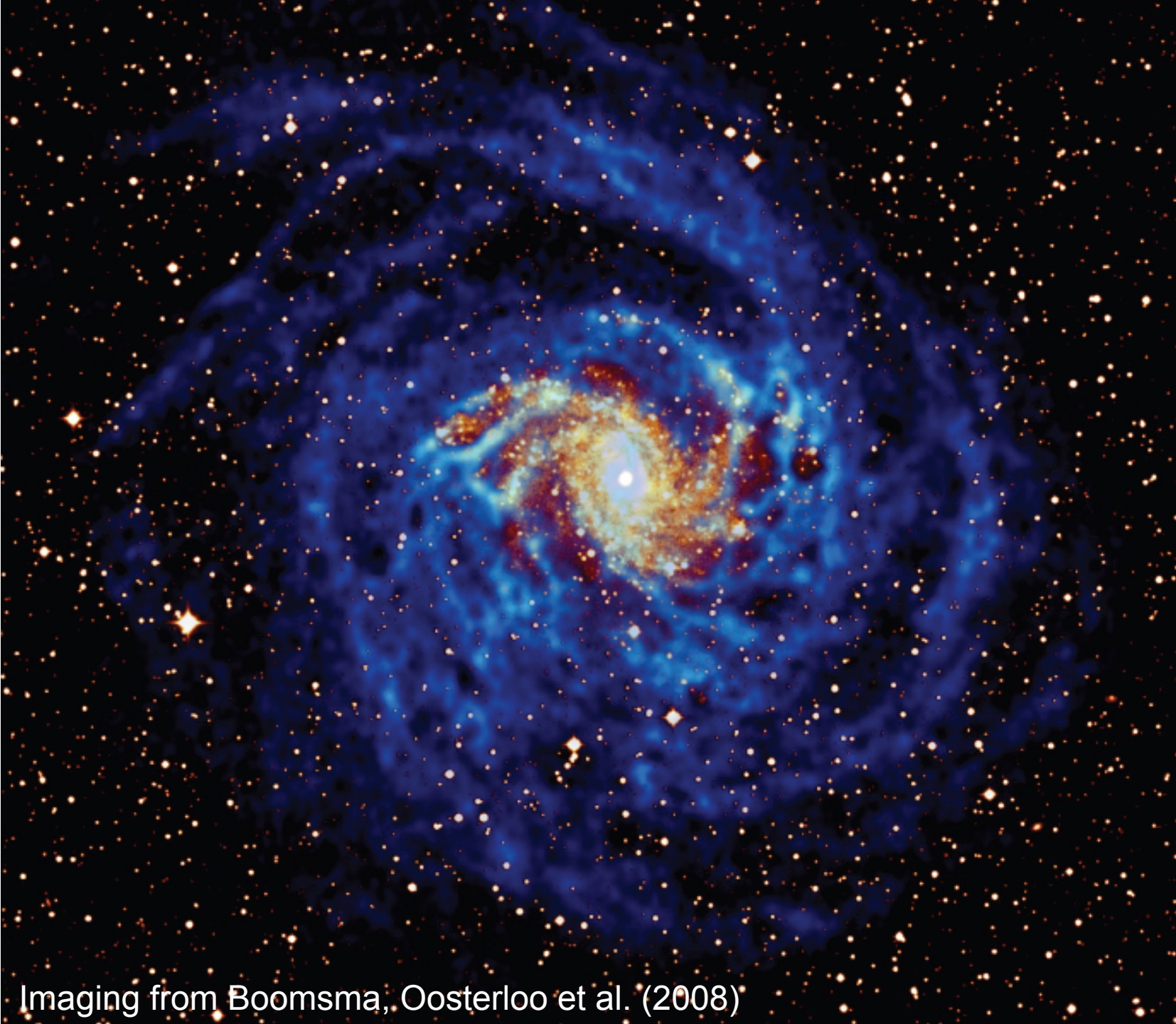
WSRT Imaging of NGC6946

16x12h with WSRT

0.2 mJy/beam (~ 4 km/s)

66 samples per channel, per
polarisation, per integration.

MeerKAT?



The benchmark

WSRT Imaging of NGC6946

16x12h with WSRT

0.2 mJy/beam (~ 4 km/s)

66 samples per channel, per
polarisation, per integration.

MeerKAT can achieve this in 12h

0.2 mJy/beam (5km/s)

2016 samples per channel, per
polarisation, per integration.



Science

Imaging

Data
Management

Storage

€

The benchmark

WSRT Imaging of NGC6946

16x12h with WSRT

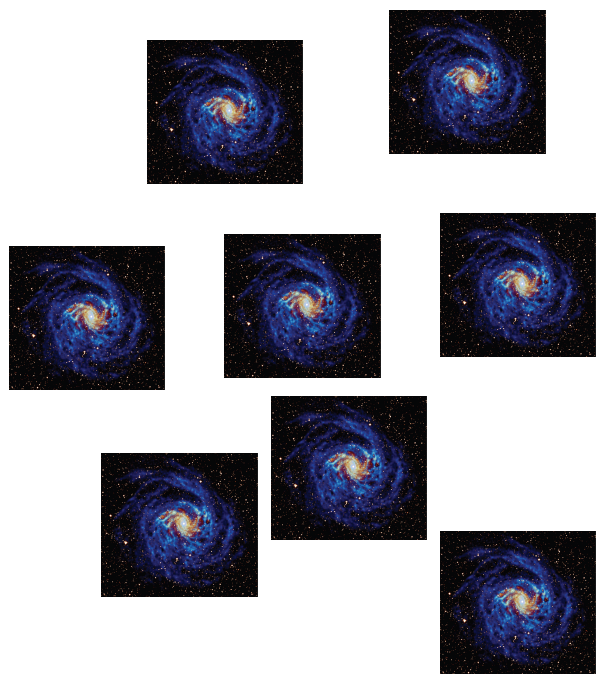
0.2 mJy/beam (~ 4 km/s)

66 samples per channel, per
polarisation, per integration.

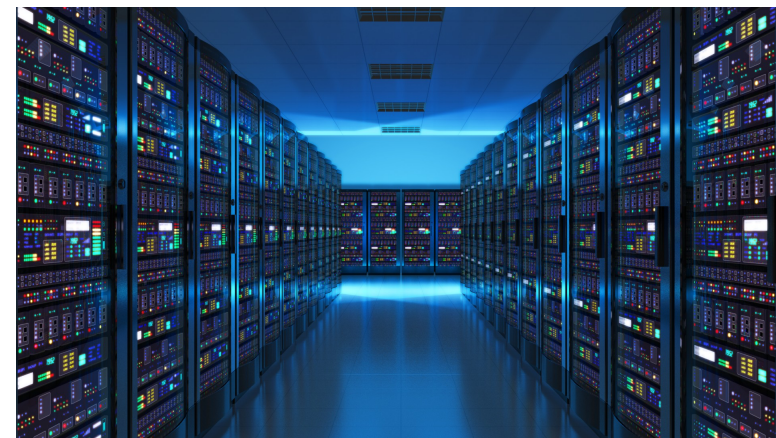
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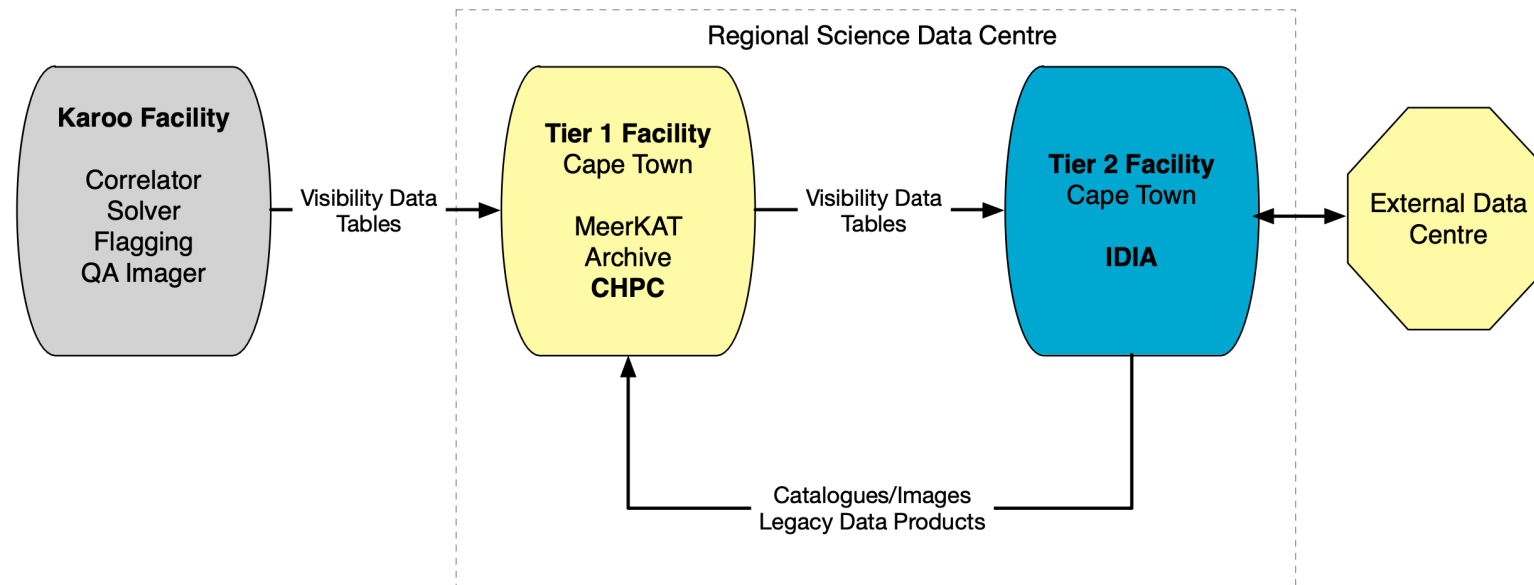
Science!



Bricks and mortar
People
Money

MIGHTEE-HI Data Processing

- Volumes Expected: 1TB/hour
- Science Processing: data transport, management and imaging
- Analysis: Interactive, collaborative, agile, user-focused



Data Processing for the SKA*

- Not entirely unexpected...



Square Kilometre Array
Software and Computing:

Estimating the Sizes of the Challenges

Duncan Hall

SPDO Software and Computing

2009 Oct 13



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* and SKA pathfinders

Data Processing for the SKA*

Article

Challenges in exascale radio astronomy: Can the SKA ride the technology wave?

Erik Vermij¹, Leandro Fiorin¹, Rik Jongerius¹, Christoph Hagleitner² and Koen Bertels³

Abstract

The Square Kilometre Array (SKA) will be the most sensitive radio telescope in the world. This unprecedented sensitivity will be achieved by combining and analyzing signals from 262,144 antennas and 350 dishes at a raw data rate of petabits per second. The processing pipeline to create useful astronomical data will require hundreds of peta-operations per second, at a very limited power budget. We analyze the compute, memory and bandwidth requirements for the key algorithms used in the SKA. By studying their implementation on existing platforms, we show that most algorithms have properties that map inefficiently on current hardware, such as a low compute–bandwidth ratio and complex arithmetic. In addition, we estimate the power breakdown on CPUs and GPUs, analyze the cache behavior on CPUs, and discuss possible improvements. This work is complemented with an analysis of supercomputer trends, which demonstrates that current efforts to use commercial off-the-shelf accelerators results in a two to three times smaller improvement in compute capabilities and power efficiency than custom built machines. **We conclude that waiting for new technology to arrive will not give us the instruments currently planned in 2018: one or two orders of magnitude better power efficiency and compute capabilities are required. Novel hardware and system architectures, to match the needs and features of this unique project, must be developed.**

International Journal of
HIGH PERFORMANCE
COMPUTING APPLICATIONS

The International Journal of High
Performance Computing Applications
2015, Vol. 29(1) 37–50
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DOI: 10.1177/1094342014549059
hpc.sagepub.com


- ~2015
- Current state of the art, and projected trends, are *insufficient for the exascale challenge of the SKA.*

* and SKA pathfinders

Data Processing for the SKA*

Article

International Journal of
HIGH PERFORMANCE
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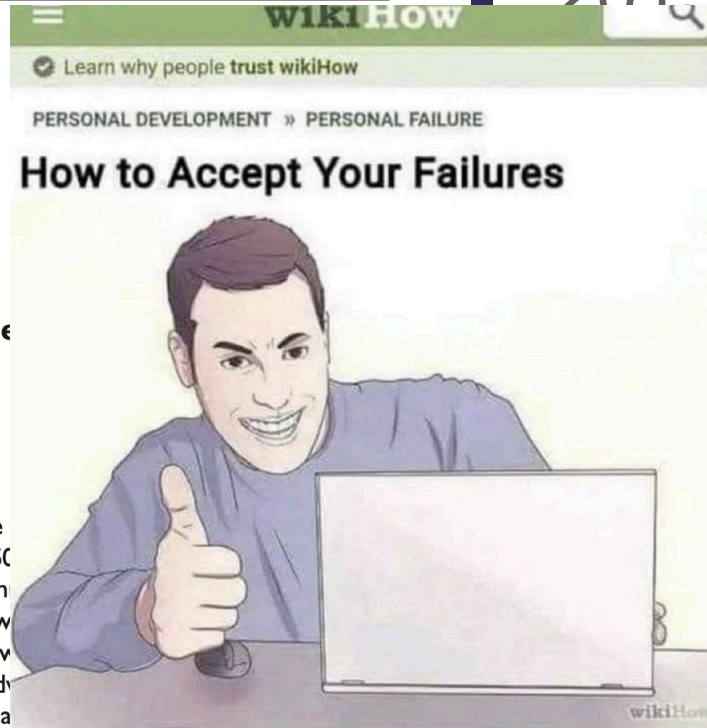
~2015

Challenges in exascale radio astronomy: Can the SKA ride the technology wave?

Erik Vermij¹, Leandro Fiorin¹, Rik Jongerius¹, Christoph Hagler² and Koen Bertels³

Abstract

The Square Kilometre Array (SKA) will be the most sensitive radio telescope in the world. It will be achieved by combining and analyzing signals from 262,144 antennas and 350 TeraFLOPS per second. The processing pipeline to create useful astronomical data will require high performance computing, at a very limited power budget. We analyze the compute, memory and bandwidth requirements used in the SKA. By studying their implementation on existing platforms, we identify properties that map inefficiently on current hardware, such as a low compute-bandwidth ratio. In addition, we estimate the power breakdown on CPUs and GPUs, analyze the capabilities of current efforts to use commercial off-the-shelf accelerators results in a two to three orders of magnitude better power efficiency and compute capabilities than custom built machines. We conclude that the hardware that will arrive will not give us the instruments currently planned in 2018: one or two orders of magnitude better power efficiency and compute capabilities are required. Novel hardware and system architectures, to match the needs and features of this unique project, must be developed.



1 Remember one day the sun will explode.
and no one will even be alive to remember you.

It state of the art, projected trends, are sufficient for the scale challenge of the

* and SKA pathfinders

Data Processing for the SKA*



Computing Challenges in the SKA Era

Bruce Elmegreen

IBM T.J. Watson Research Center

Yorktown Heights, NY 10598

bge@us.ibm.com

Rutgers University, March 16-18, 2015

Summary

- Processors: Moore's law for systems still on track
 - ** \$T/year investment: exponential growth continues to the SKA1 era
 - ** also assumed by the SKA SDP 2015 report
- Memory: changing to become more energy efficient (non-volatile)
 - also allows extremely fast large memory spaces (“solid state memory”)
 - However, new memory technologies are more expensive now
- SKA1 data volumes are not excessive by world standards
- SKA1 raw data rates are large but the technology should handle it.
- Not discussed: software changes, machine learning, neural networks, ...

* and SKA pathfinders

Data Processing for the SKA*

R.T. Schilizzi et al. The Square Kilometre Array - an overview

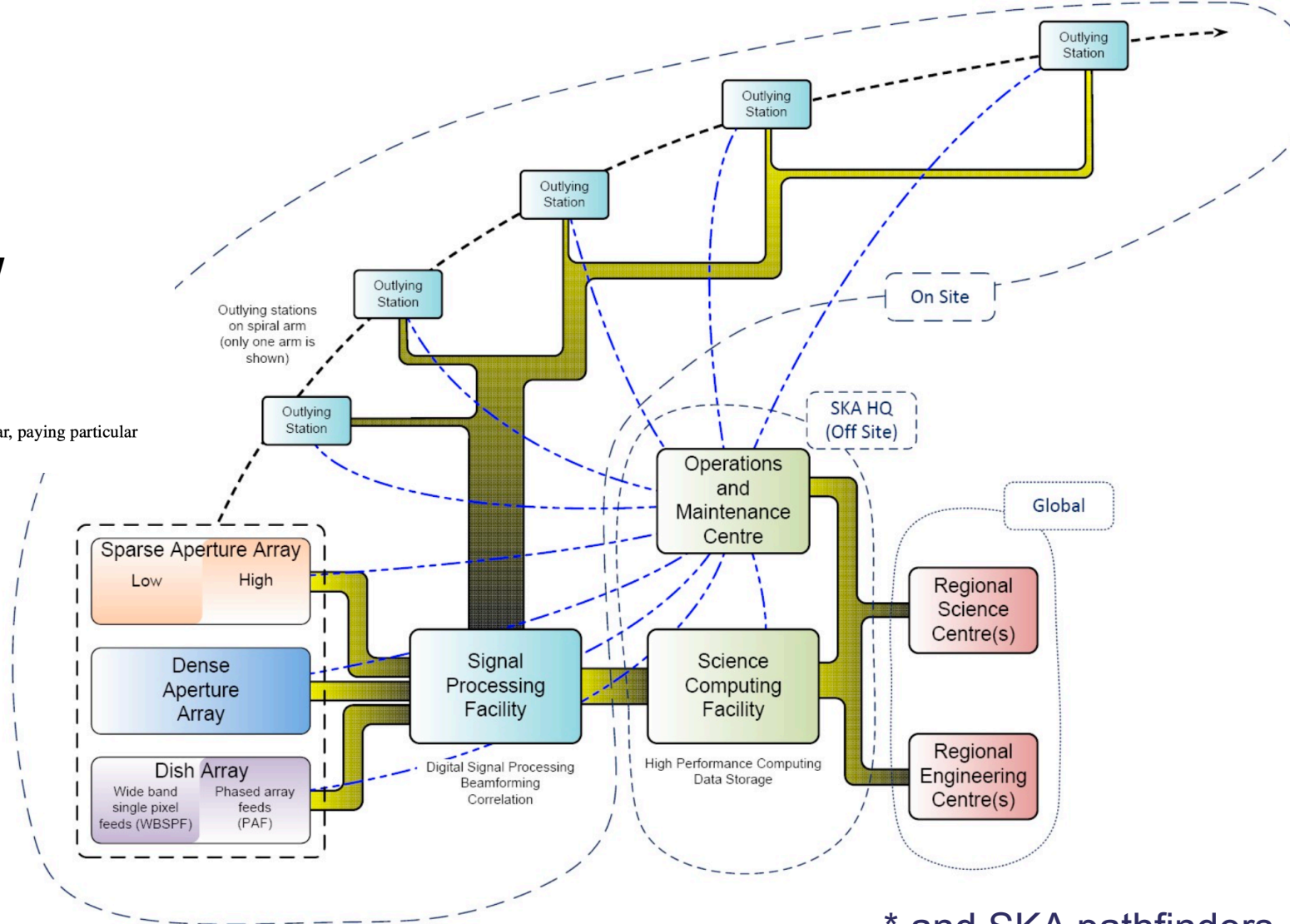
WIDEFIELD SCIENCE AND TECHNOLOGY FOR THE SKA
 SKADS CONFERENCE 2009
 S.A. Torchinsky, A. van Ardenne, T. van den Brink-Havinga, A.J.J. van Es, A.J. Faulkner (eds.)
 4-6 November 2009, Château de Limelette, Belgium

The Square Kilometre Array - an overview

R.T. Schilizzi¹, P.E. Dewdney¹, and T.J.W. Lazio²

¹ SKA Program Development Office, University of Manchester, UK
² Naval Research Laboratory, Washington DC, USA

Abstract. A brief overview is given of the international SKA project and developments during the past year, paying particular attention to technical highlights, site characterisation, schedule and policy issues.



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Data Processing for the SKA*

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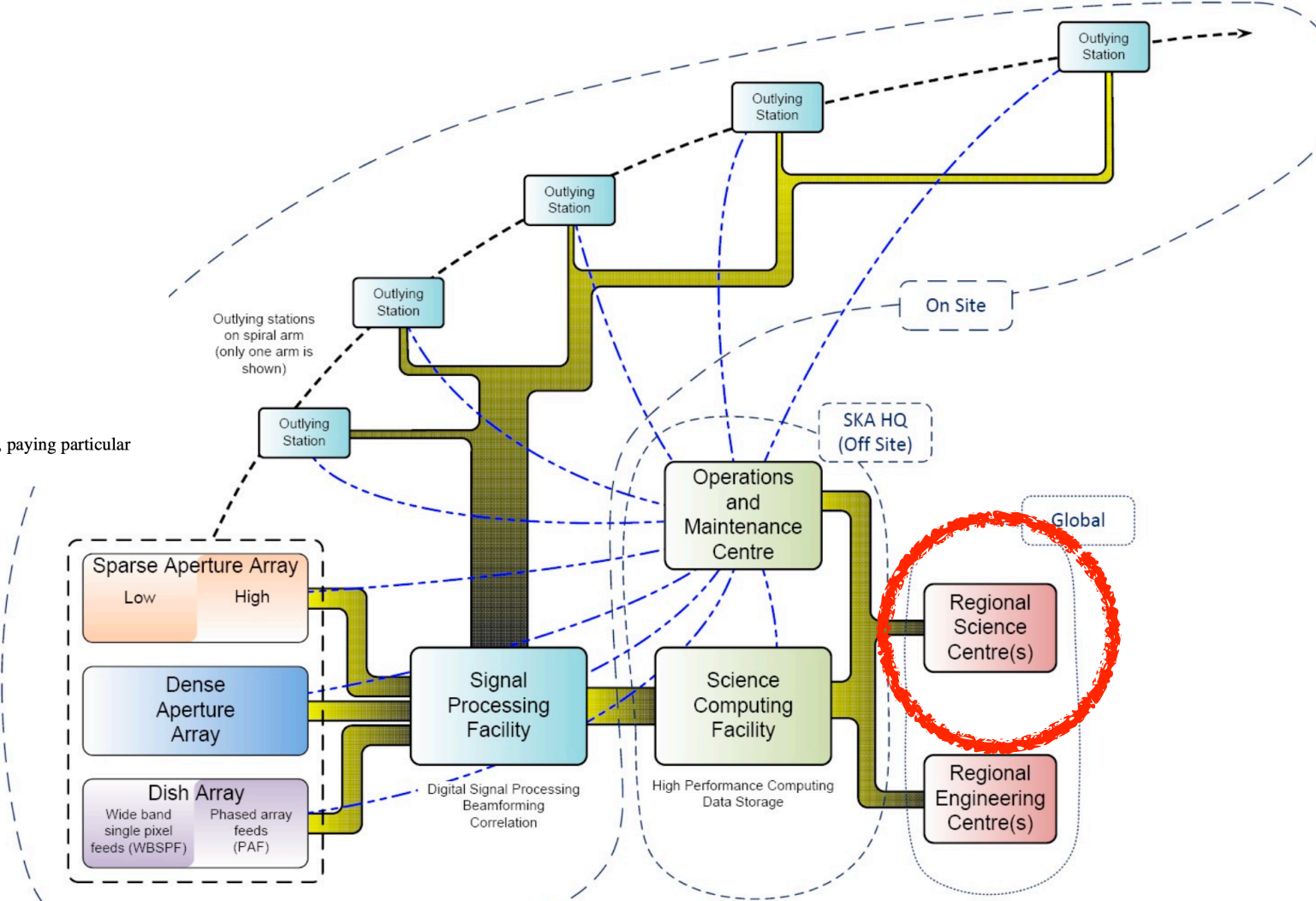
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Home > News > Speeches > Minister Naledi Pandor: Launch of Inter-University Instit...

Keyword

Categories

Speeches

Government Level

- Any -

Province

- Any -

Subjects

- 16 days of activism
- 20 years of freedom
- Africa
- African Peer Review Mechanism
- Agriculture
- Anti-corruption initiatives
- Arts and culture
- Aviation
- Black Economic Empowerment

Start Date

09/09/2024

Minister Naledi Pandor: Launch of Inter-University Institute for Data Intensive Astronomy (IDIA)

03 Sep 2015

Address by Naledi Pandor MP, Minister of Science and Technology, at the launch of the Inter-University Institute for Data Intensive Astronomy (IDIA), South African Astronomical Observatory, Cape Town

Prof Russ Taylor
 Prof Tyrone Pretorius, UWC Vice-Chancellor,
 Dr Max Price, UCT Vice-Chancellor,
 Dr Bernie Fanaroff
 Prof Frik van Niekerk, NWU Deputy Vice-Chancellor,
 Directors of the SKA project



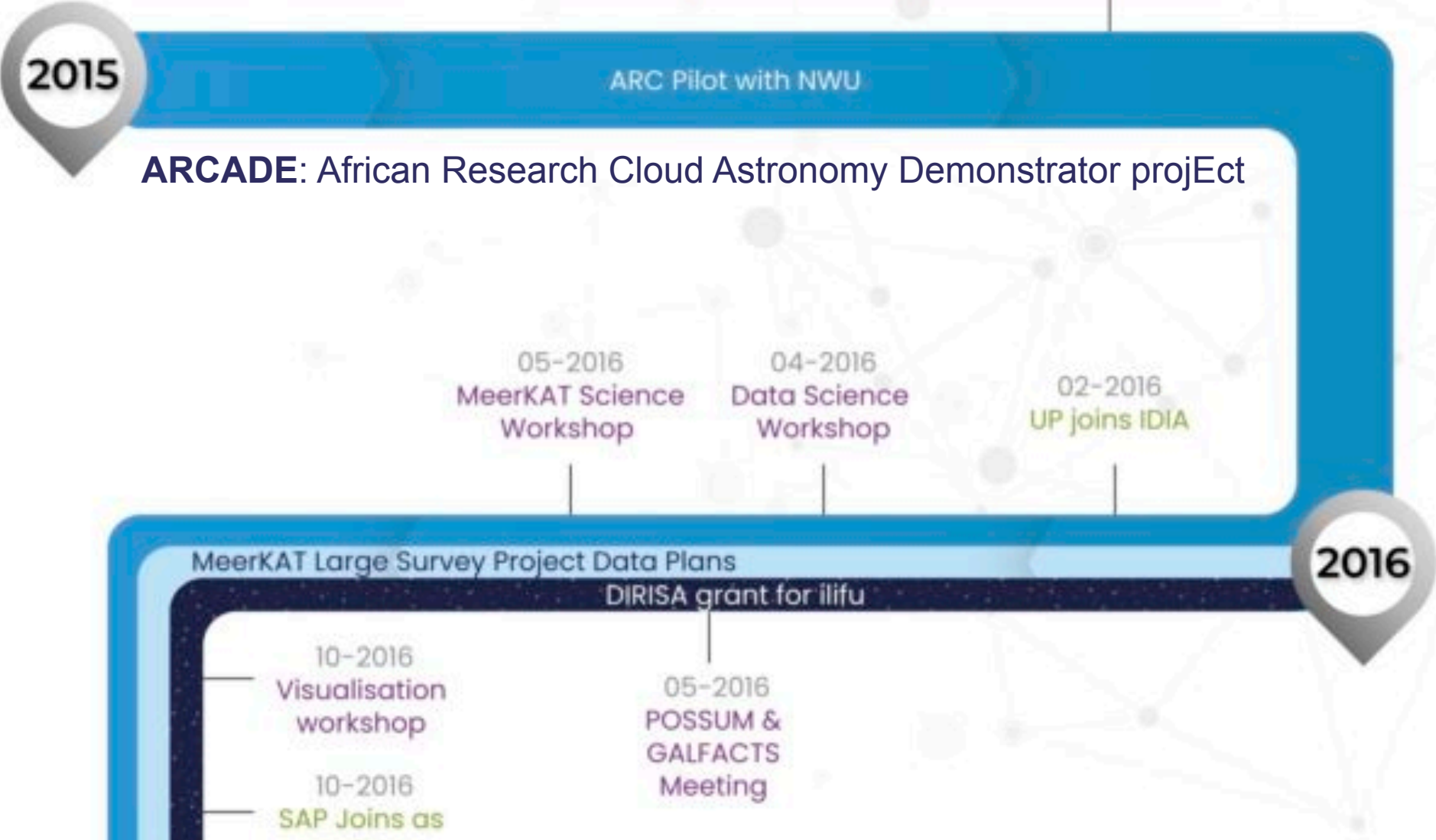
IDIA

- Facility to empower the science users.
- Quid pro quo — scientific community would need to *show up* (\$).
- Partnership with observatory, but community driven.
- Political and scientific support, and \$ to get started.

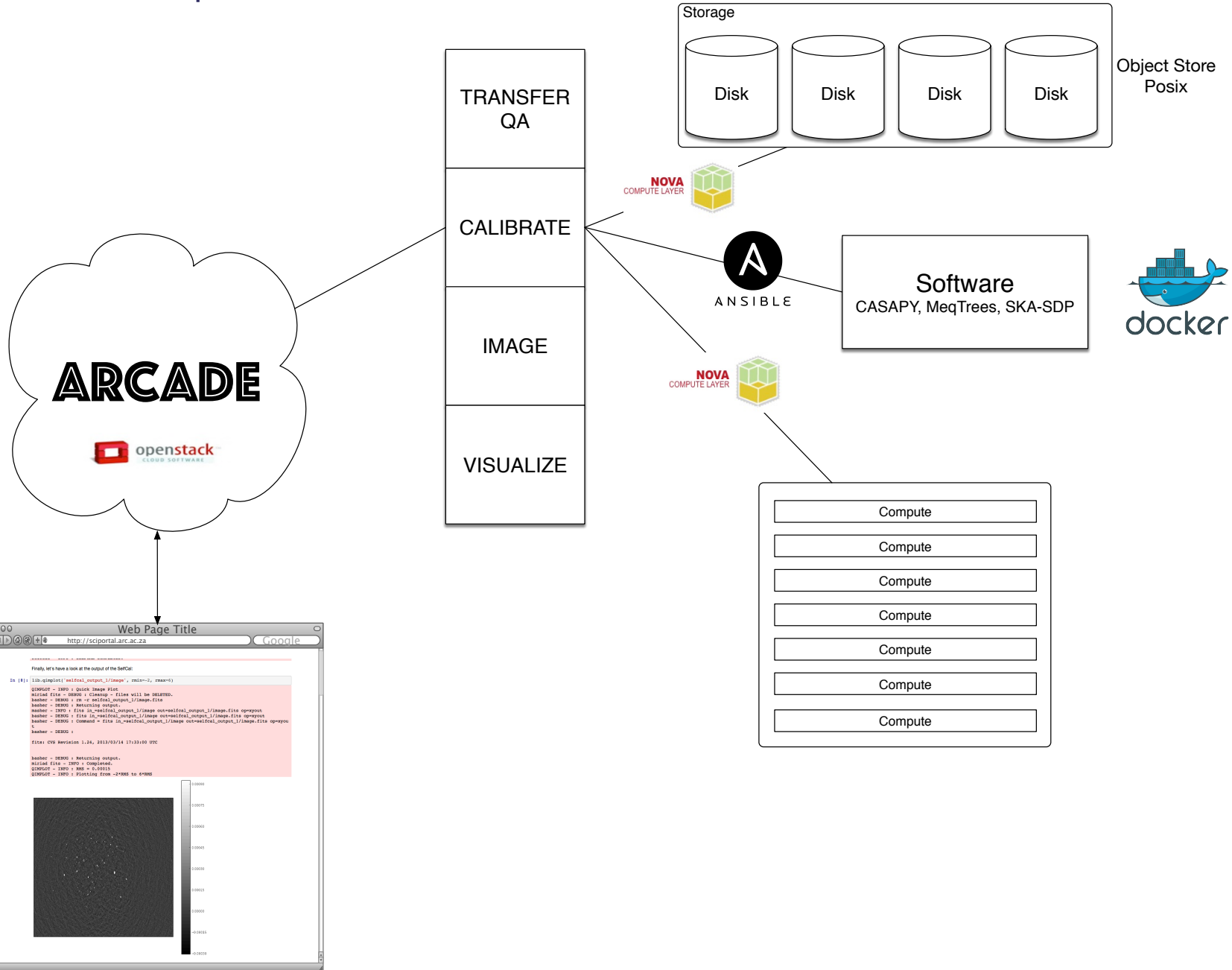
IDIA

- Facility to empower the science users.
- Quid pro quo — scientific community would need to show up (\$).
- Partnership with observatory, but community driven.
- Political and scientific support, and \$ to get started.
- But therein lay the problem — how do we get started?

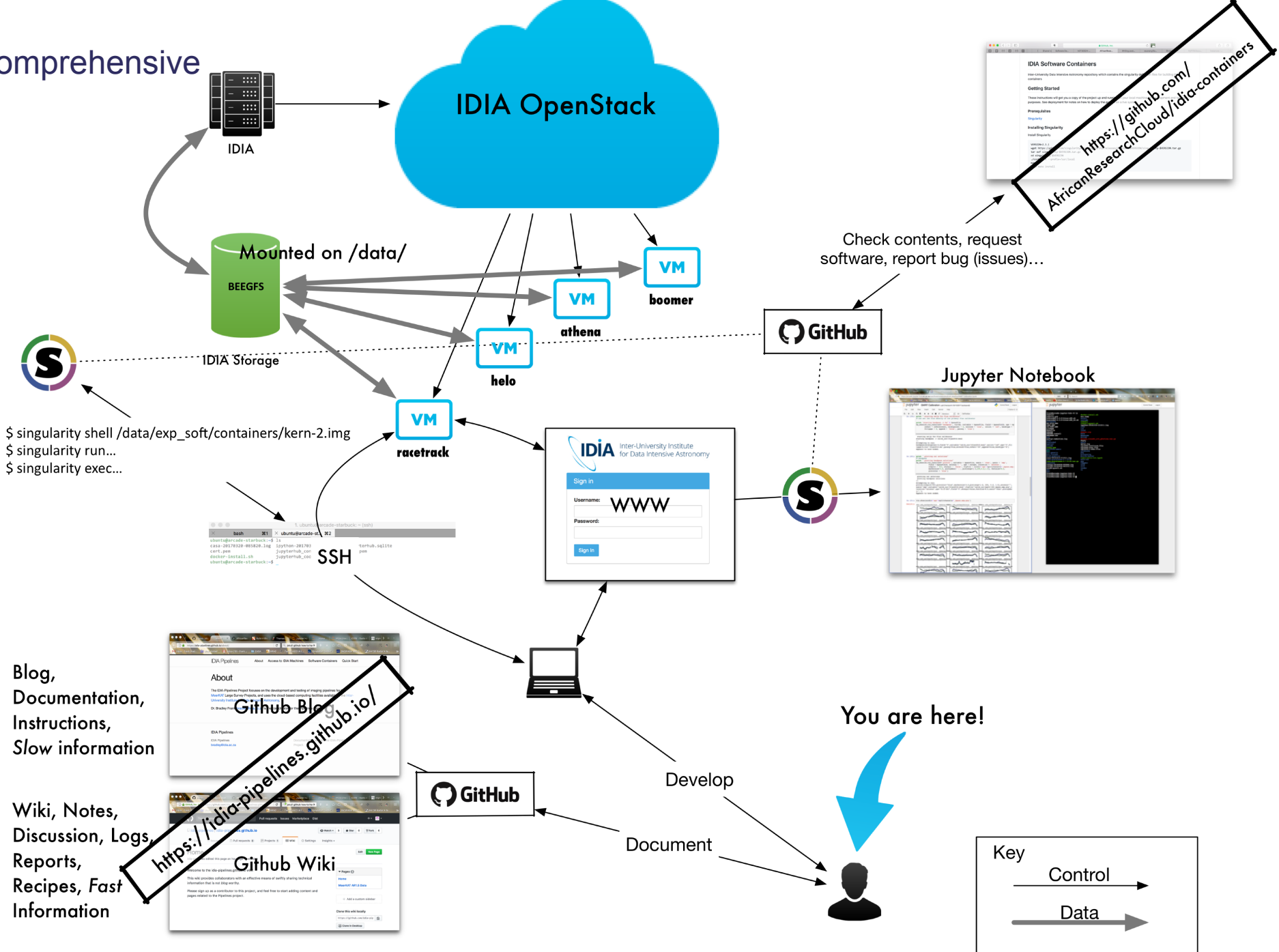
IDIA Timeline



Basic expression of scientific requirements...



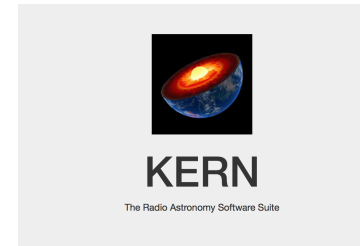
Expanded into a comprehensive architecture...



Further broken down into component design...

```

1 {
2   "display_name": "KERN-2",
3   "language": "python",
4   "argv": [
5     "/data/exp_soft/containers/kern-2.img",
6     "python",
7     "-m",
8     "ipykernel_launcher",
9     "-f",
10    "{connection_file}"
11  ]
12 }
    
```



- What is KERN? +
- Quickly getting started +
- News? +
- How do I use it? +
- What packages are available? +
- Why should I use KERN? +



jupyter Tidal Forces on the Earth Last Checkpoint: 04/25/2017 (autosaved)

File Edit View Insert Cell Kernel Help Python [Root] O

Differential Force on the Earth due to the Moon

```

In [2]: import numpy as np
import pylab as pl
import matplotlib inline

/Users/frank/anaconda/lib/python2.7/site-packages/matplotlib/font_manager.py:273: UserWarning: Matplotlib is building the font cache using fc-list. This may take a moment.
warnings.warn("Matplotlib is building the font cache using fc-list. This may take a moment.")

The differential force on the Earth due to the Moon is:

$$\Delta F \approx \frac{GM_{\text{moon}}}{r^2} \sqrt{2} (\cos \theta - j \sin \theta)$$

The magnitude is:

$$|\Delta F| = k^2 \frac{M_{\text{moon}}}{r^2} \sqrt{4 \cos^2 \theta + \sin^2 \theta}$$

where  $k = GM_{\text{moon}}$  .

In the next cell I'm going to define a few important variables.

In [3]: R_earth = 6000.
r = 384400.
theta = pi.arange(-2*pi, pi, 2*pi, 0.1)

Now I'm simply going to calculate the differential force due to the Moon, considering distances of 0.4r to r. I don't consider smaller distances, since the force will be much greater and the plotting will screw up.

In [4]: r = 0.4*384400.
df = (R_earth/(r**3))*(4.*pl.cos(theta)**2+pl.sin(theta)**2)
pl.polar(theta, df, label='40%')

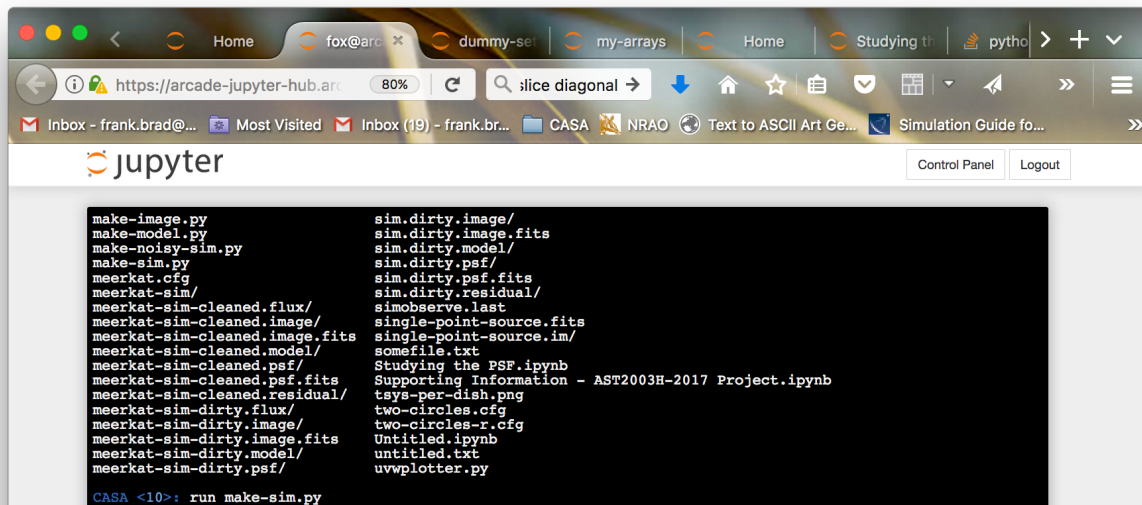
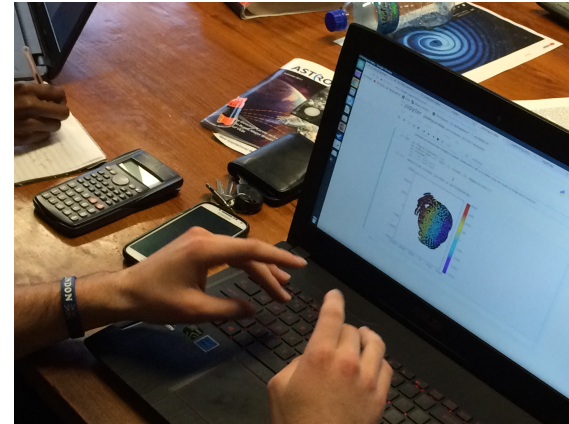
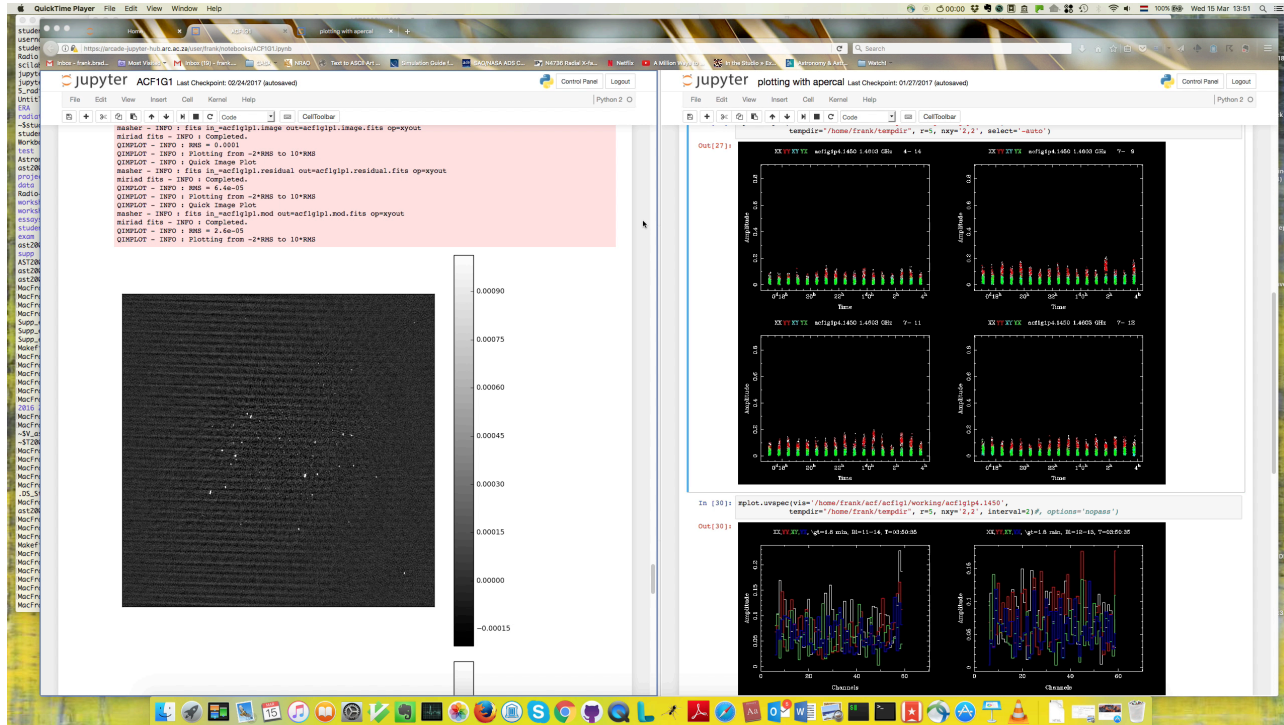
r = 0.6*384400.
df = (R_earth/(r**3))*(4.*pl.cos(theta)**2+pl.sin(theta)**2)
pl.polar(theta, df, label='60%')

r = 0.8*384400.
df = (R_earth/(r**3))*(4.*pl.cos(theta)**2+pl.sin(theta)**2)
pl.polar(theta, df, label='80%')

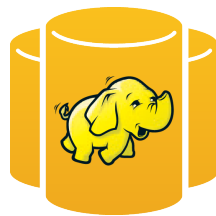
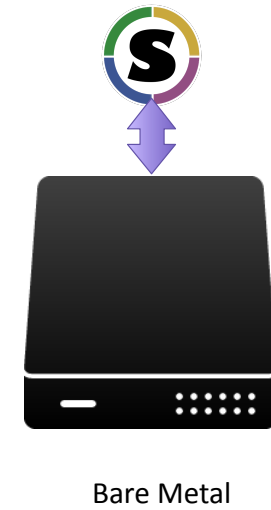
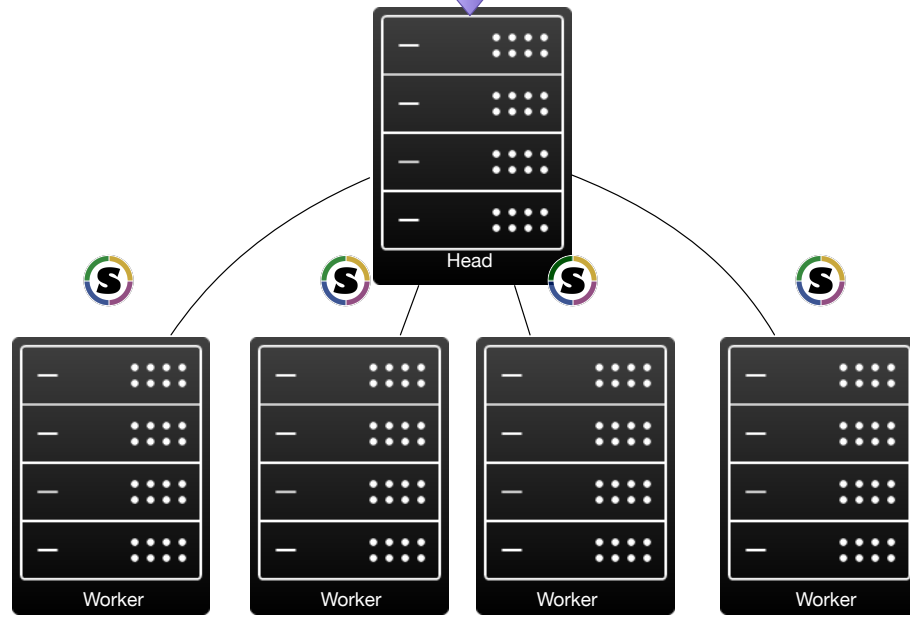
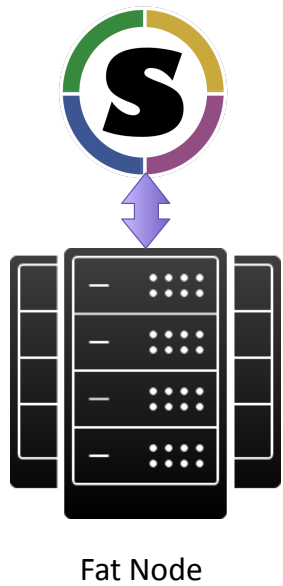
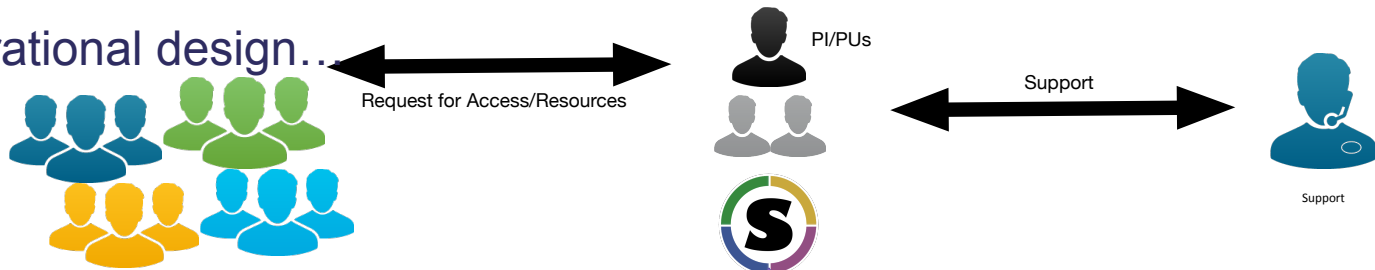
r = 1.0*384400.
df = (R_earth/(r**3))*(4.*pl.cos(theta)**2+pl.sin(theta)**2)
pl.polar(theta, df, label='100%')
pl.legend()

Out[4]: <matplotlib.legend.Legend at 0x10cbe4390>
    
```

Substantiated by scientific use-cases...



Scaled to an operational design...



Hadoop/Distributed Storage



Posix

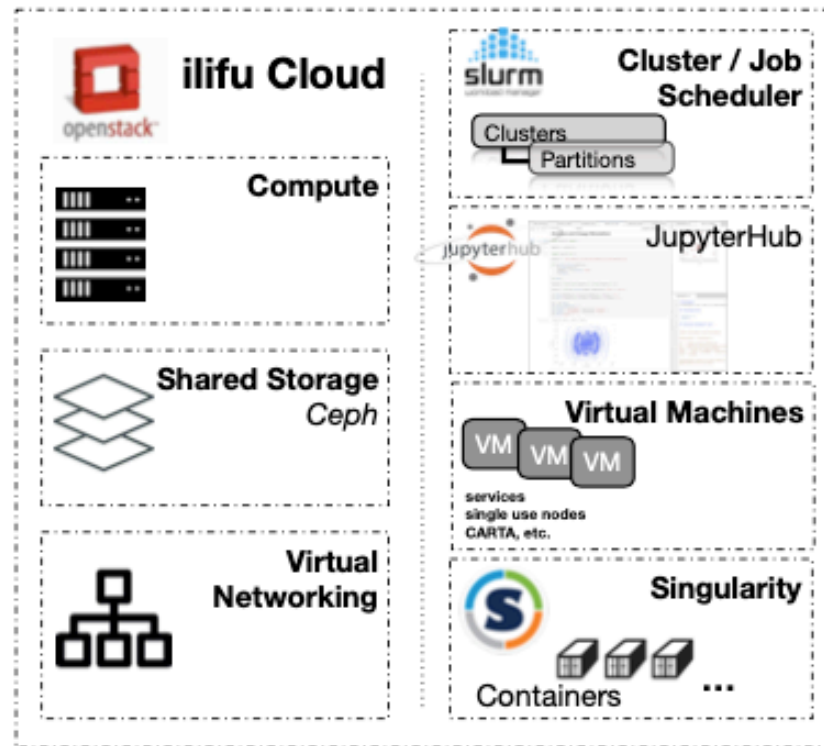
Recap (2015-2017)

- Used the African Research Cloud as a pathfinder for IDIA design.
- Defined the science requirements.
- Data processing use-cases: WSRT data, MeerKAT-16.
- Analysis: large astronomy groups, distributed collaborators...
- Laid the technical and technological groundwork for IDIA development, Ilifu procurement and data centre design.
- **Convinced ourselves that we could build complex systems.**

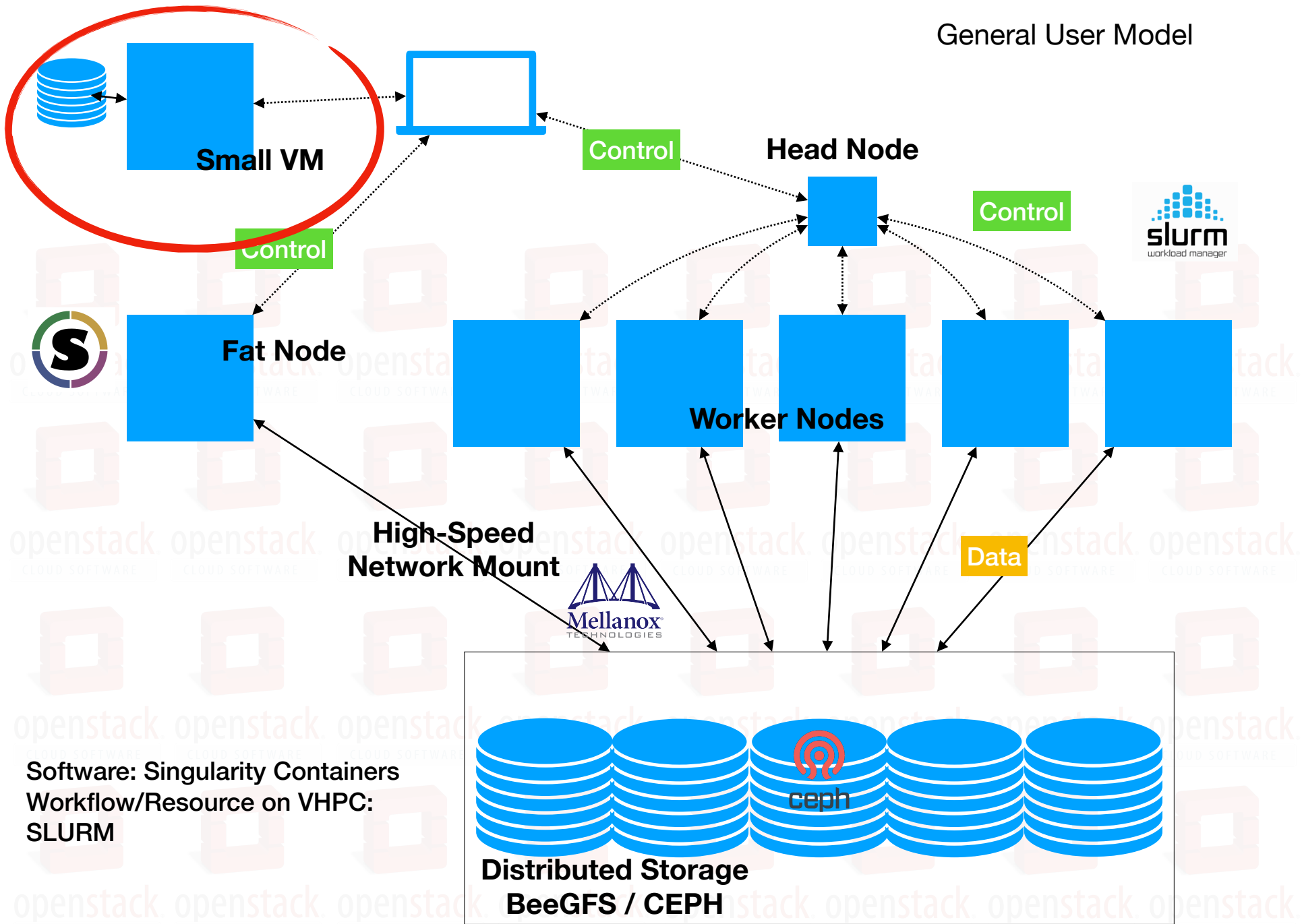
IDIA and Ilifu (2017+)

- Partnership with Bioinformatics stakeholders.
- Included the observatory (SARAO) as a partner.

ilifu Computing Environment



General User Model



Software: Singularity Containers
 Workflow/Resource on VHPC:
 SLURM

Distributed Storage
BeeGFS / CEPH

32KN + IDIA →

Erwin de Blok
SCI-20180516-EB-01
20210308-0066
1647545027

J0408-6545
0408-658
J1130-1449
1127-145
...

MHONGOOSE J1337-28 rising 4
1336.499999832362 MHz to 1443.499999832362 MHz
Run at 2022-03-17 19:25:22 UTC



Details

[Script Log](#)

Delay Cal Observation

[Calibration Report 1](#) [Calibration Report 2](#) [Calibration Report 3](#) [Calibration Report 4](#)

4KW + IDIA →

Erwin de Blok
SCI-20180516-EB-01
20210308-0066
1647545021

J0408-6545
0408-658
J1130-1449
1127-145
...

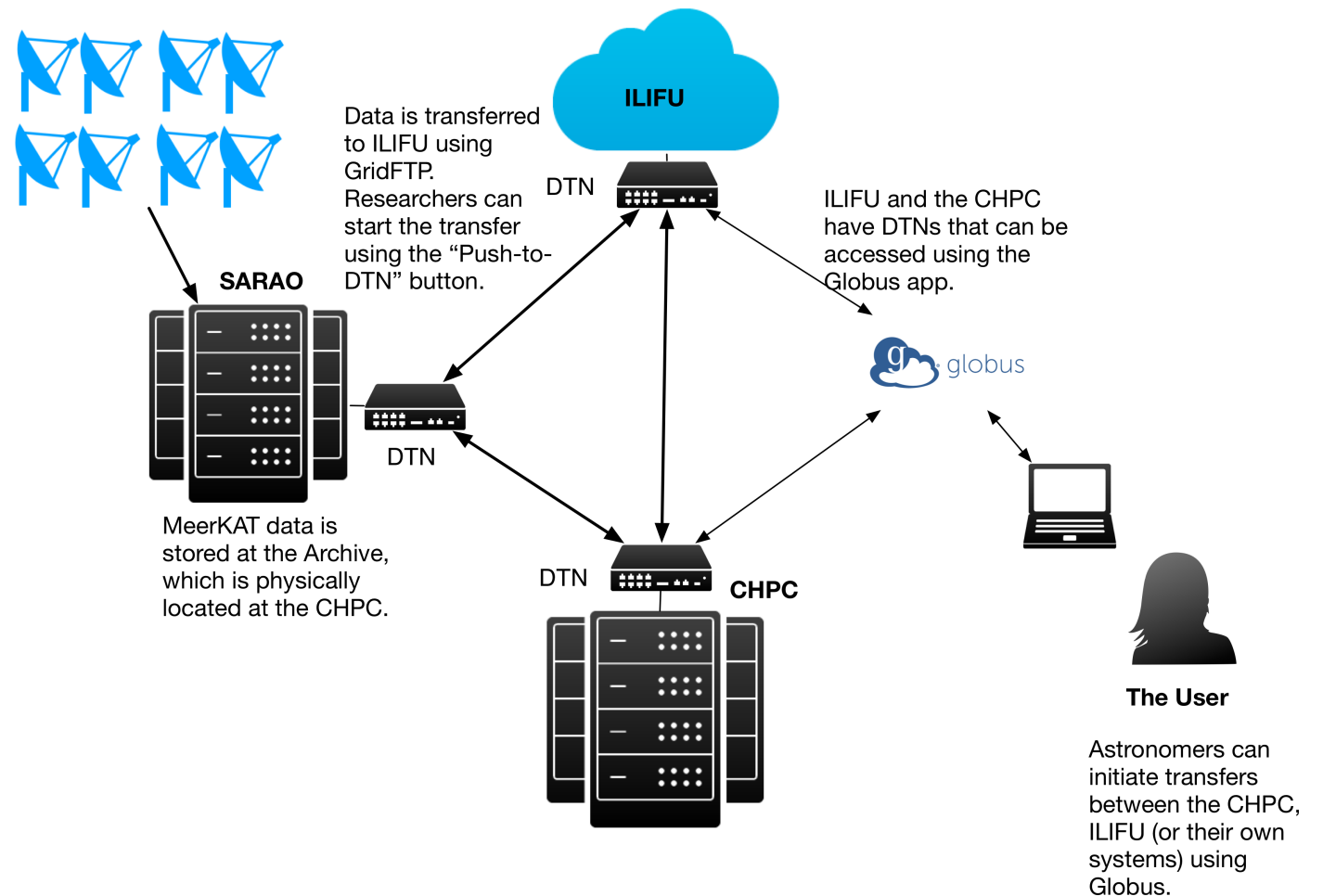
MHONGOOSE J1337-28 rising 4
856 MHz to 1712 MHz
Run at 2022-03-17 19:25:14 UTC

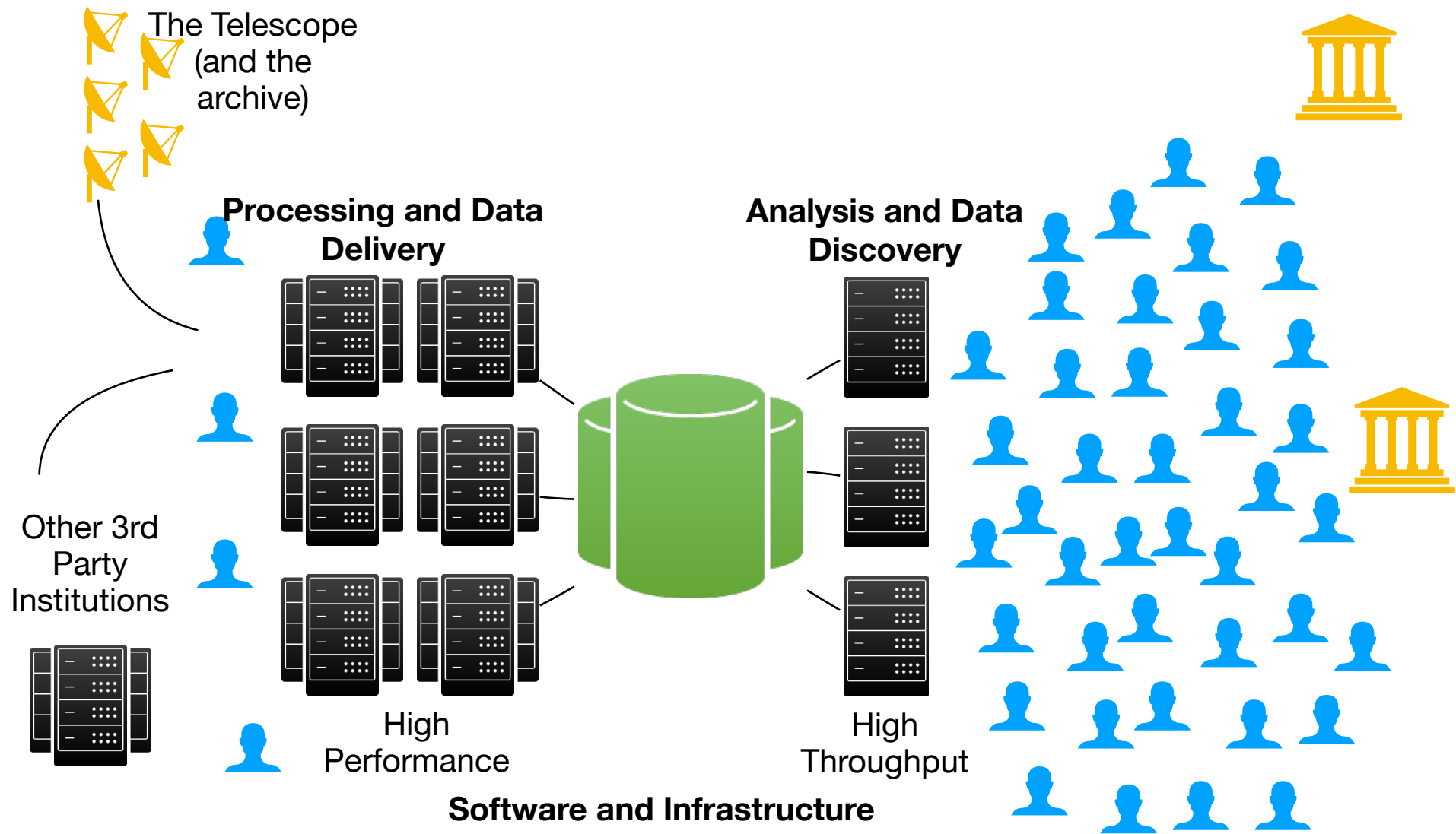
Details

[Script Log](#)

Delay Cal Observation

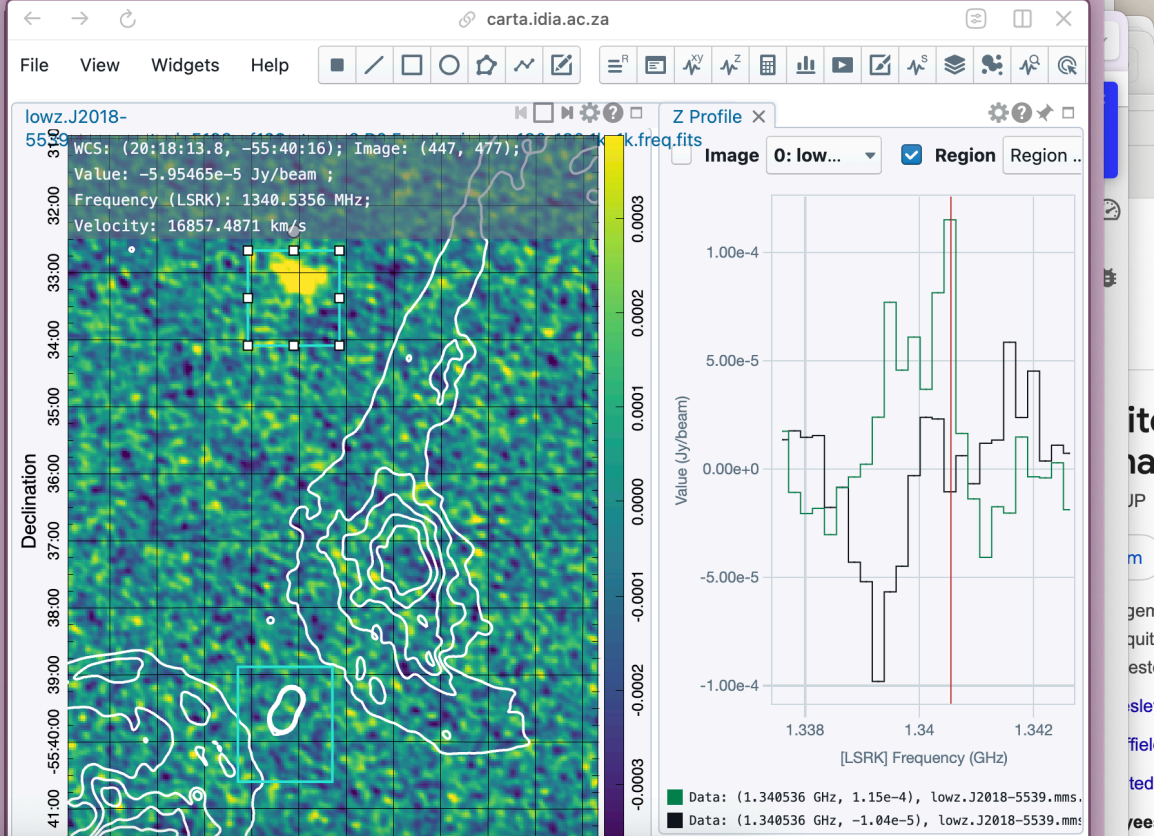
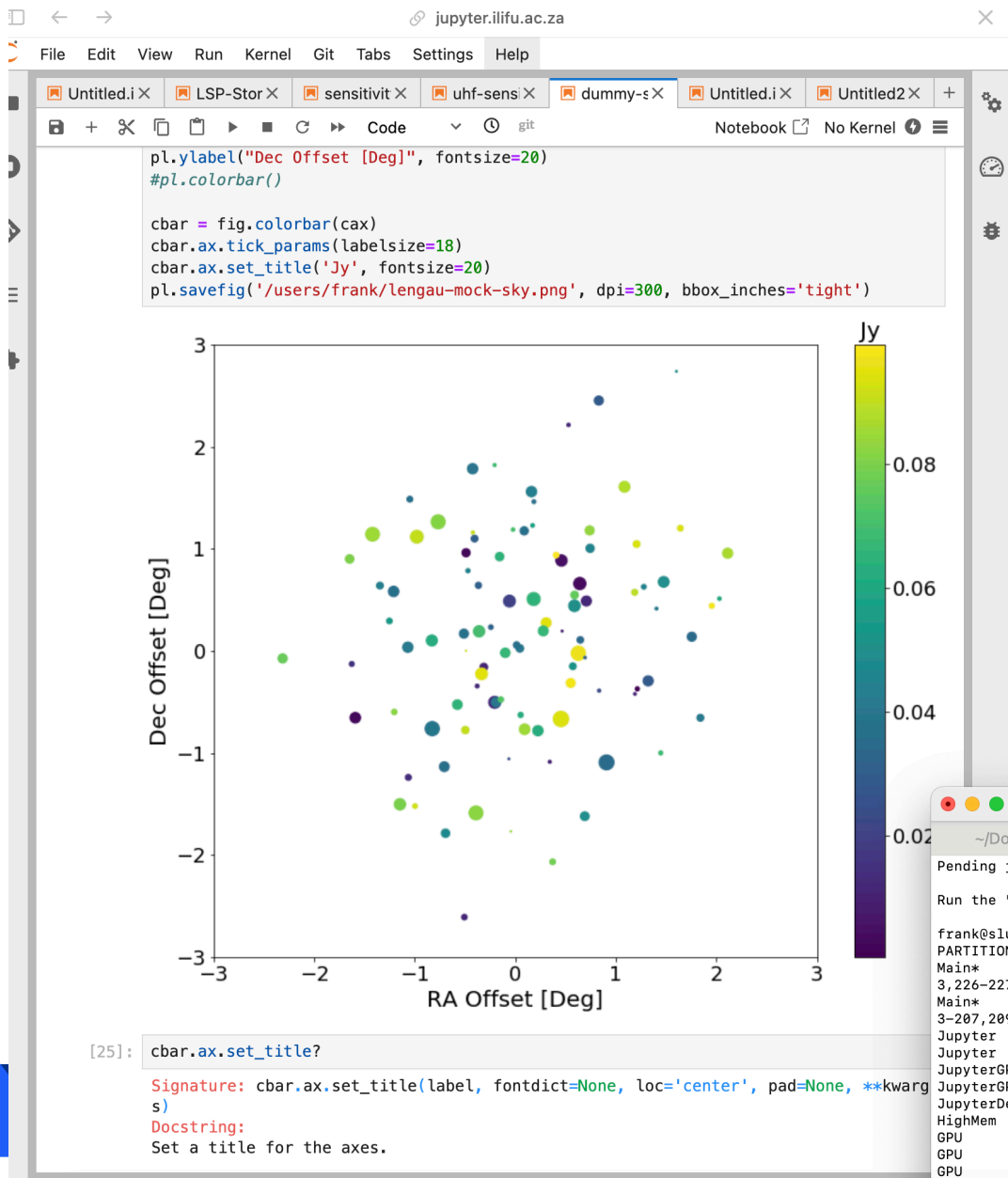
- Design emerged after coordination/ collaboration with SRAO, CHPC and IDIA.
- Hackathons/documents/designs/ negotiation.
- SANReN port + redirection.
- Staging (raw->MS), transfer, authentication and then validation on the IDIA side.
- Security — authenticate on the SRAO side to *request* a push, validated identity, submitted to queue and then transferred back using certificates.





- A few users, long compute times, large data volumes.
- Changes slowly.
- Changes are big. Affect both software *and* infrastructure.

- Many users, highly interactive. Smaller data volumes.
- Changes often, but changes are small.



```

Downloads — ssh frank@slurm.ilifu.ac.za -i ilifu -Y — 93x21
~/Downloads — jupyter-notebook > python ...s — ssh frank@slurm.ilifu.ac.za -i ilifu -Y
Pending job count: 0

Run the "shelp" command to display this message.

frank@slurm-login:~$ sinfo
PARTITION AVAIL TIMELIMIT NODES STATE NODELIST
Main* up 14-00:00:0 27 mix compute-[002,005-006,008,013,016,101-105,202,208,223,226-227,229-230,233,235,237,240-242,248-250]
Main* up 14-00:00:0 55 alloc compute-[003-004,007,009-012,014-015,017-018,201,203-207,209-222,224-225,228,231-232,234,236,238-239,243-247,251-260]
Jupyter up infinite 4 mix jupyter-[005,009-010,013]
Jupyter up infinite 8 alloc jupyter-[002-004,006-008,011-012]
JupyterGPU up 14-00:00:0 1 mix gpu-003
JupyterGPU up 14-00:00:0 1 idle gpu-004
JupyterDev up infinite 1 alloc jupyter-001
HighMem up 14-00:00:0 3 mix highmem-[001-003]
GPU up 14-00:00:0 2 mix gpu-[003,007]
GPU up 14-00:00:0 1 alloc gpu-005
GPU up 14-00:00:0 4 idle gpu-[001-002,004,006]
Devel up 5-00:00:0 1 alloc compute-001
frank@slurm-login:~$

```



Science Successes

- 1 2022MNRAS.512.2697R 2022/05   
MIGTEE-H I: the H I size-mass relation over the last billion years
 Rajohnson, Sambatriniaina H. A.; Frank, Bradley S.; Ponomareva, Anastasia A. *and 15 more*
- 2 2022MNRAS.tmp..971T 2022/04   
MIGTEE-H I: The relation between the H I gas in galaxies and the cosmic web
 Tudorache, Madalina N.; Jarvis, M. J.; Heywood, I. *and 11 more*
- 3 2022MNRAS.509.2150H 2022/01 *cited: 5*   
MIGTEE: total intensity radio continuum imaging and the COSMOS/XMM-LSS Early Science fields
 Heywood, I.; Jarvis, M. J.; Hale, C. L. *and 23 more*
- 4 2021MNRAS.508.1897P 2021/12 *cited: 1*   
Measuring the baryonic Tully-Fisher relation below the detection threshold
 Pan, Hengxing; Jarvis, Matt J.; Ponomareva, Anastasia A. *and 4 more*
- 5 2021MNRAS.508.1195P 2021/11 *cited: 4*   
MIGTEE-H I: the baryonic Tully-Fisher relation over the last billion years
 Ponomareva, Anastasia A.; Mulaudzi, Wanga; Maddox, Natasha *and 17 more*
- 6 2021MNRAS.506.2753R 2021/09 *cited: 2*   
MIGTEE-HI: discovery of an H I-rich galaxy group at $z = 0.044$ with MeerKAT
 Ranchod, Shilpa; Deane, Roger P.; Ponomareva, Anastasia A. *and 13 more*
- 7 2021MNRAS.505.2039P 2021/08 *cited: 3*   
H I intensity mapping with the MIGTEE survey: power spectrum estimates
 Paul, Sourabh; Santos, Mario G.; Townsend, Junaid *and 5 more*
- 8 2021A&A...646A..35M 2021/02 *cited: 18*   
MIGTEE-HI: The H I emission project of the MeerKAT MIGTEE survey
 Maddox, N.; Frank, B. S.; Ponomareva, A. A. *and 19 more*
- 9 2020MNRAS.491.1227P 2020/01 *cited: 4*   
Measuring the HI mass function below the detection threshold
 Pan, Hengxing; Jarvis, Matt J.; Allison, James R. *and 5 more*



UK Astronomy
Technology Centre

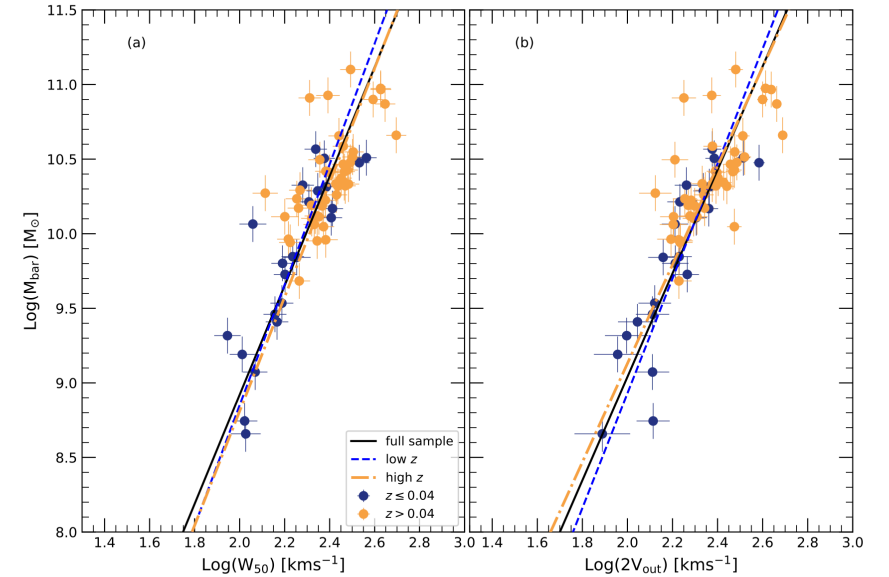
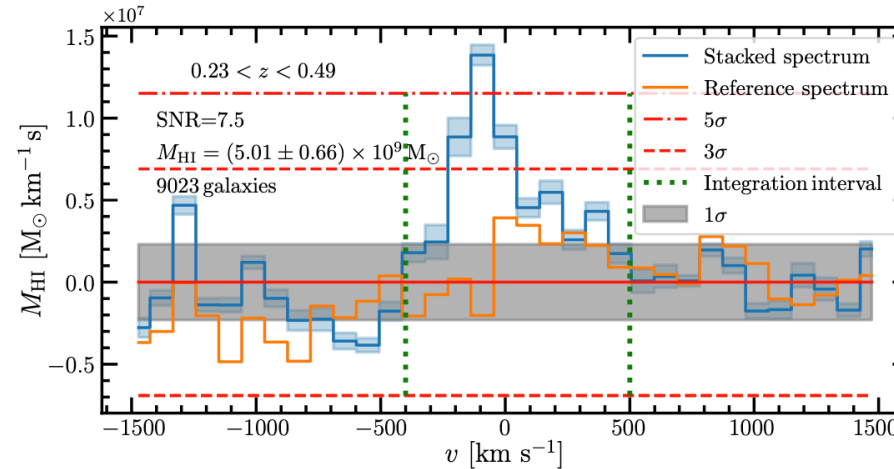
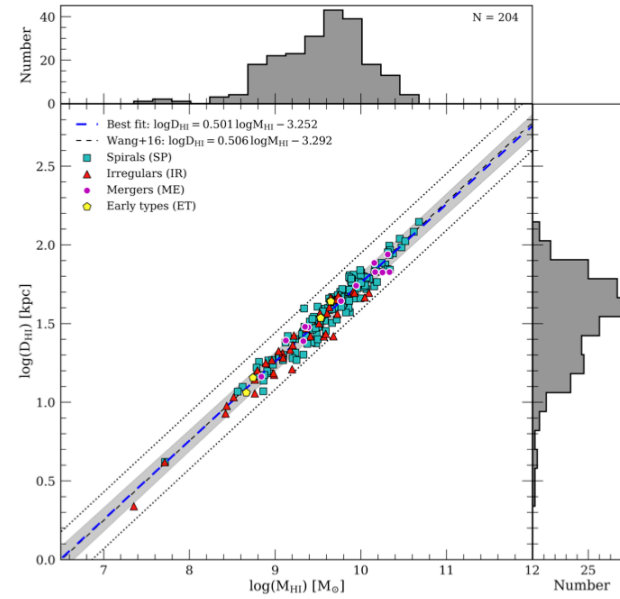


Figure 6. The bTFR based on corrected W_{50} , panel (a) and V_{out} , panel (b). The high-redshift galaxies ($z > 0.04$) are shown with orange symbols, while the low-redshift galaxies ($z \leq 0.04$) are shown with blue symbols. The best-fit for the full sample is shown with the straight black line, while the fits for low- and high- z samples are shown with blue dashed and orange dashed-dotted lines respectively.

Wanga Mulaudzi

Supervisor: Dr Bradley Frank

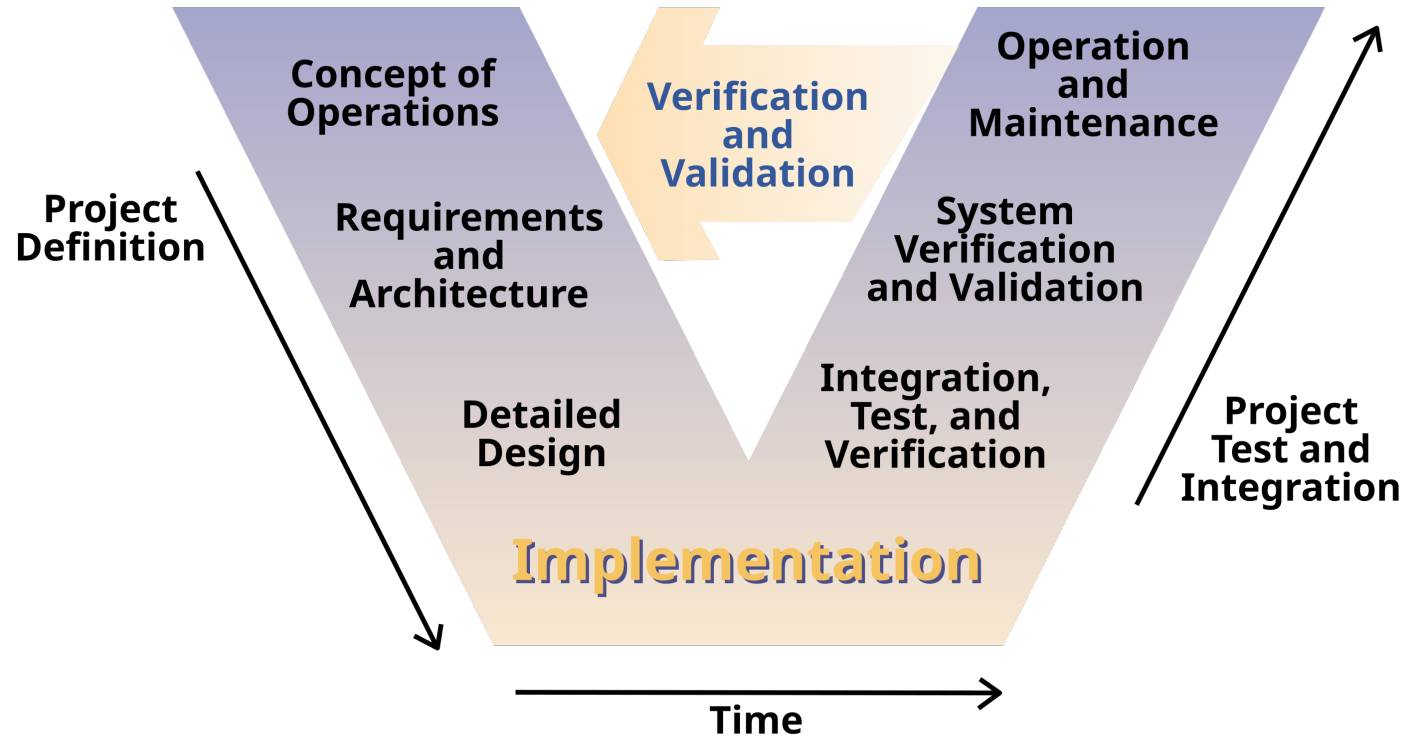
Co-Supervisor: Prof. Renée Kraan-Korteweg

Thesis presented for the degree of Master of Science



Department of Astronomy
University of Cape Town
South Africa
November 2021

Systems Engineering “V” Diagram



A&A 646, A35 (2021)

**MIGHTEE-HI: The H I emission project of the MeerKAT
 MIGHTEE survey***

N. Maddox¹, B. S. Frank^{2,3,4}, A. A. Ponomareva⁵, M. J. Jarvis^{5,6}, E. A. K. Adams^{7,8}, R. Dave^{9,6,10}, T. A. Oosterloo^{7,8}, M. G. Santos^{6,2}, S. L. Blyth⁴, M. Glowacki¹¹, R. C. Kraan-Korteweg⁴, W. Mulaudzi⁴, B. Namumba¹², I. Prandoni¹³, S. H. A. Rajohnson⁴, K. Spekkens¹⁴, N. J. Adams⁵, R. A. A. Bowler⁵, J. D. Collier^{3,15}, I. Heywood^{5,12,2}, S. Sekhar^{3,16} and A. R. Taylor^{3,11}

Received: 12 October 2020 | Accepted: 10 November 2020

Abstract
 We present the H I emission project within the MIGHTEE survey, currently being carried out with the newly commissioned MeerKAT radio telescope. This is one of the first deep, blind, medium-wide interferometric surveys for neutral hydrogen (H I) ever undertaken, extending our knowledge of H I emission to $z = 0.6$. The science goals of this medium-deep, medium-wide survey are extensive, including the evolution of the neutral gas content of galaxies over the past 5 billion years. Simulations predict nearly 3000 galaxies over $0 < z < 0.4$ will be detected directly in H I, with statistical detections extending to $z = 0.6$. The survey allows us to explore H I as a

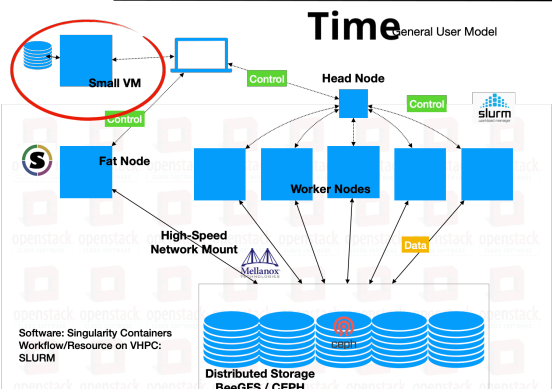
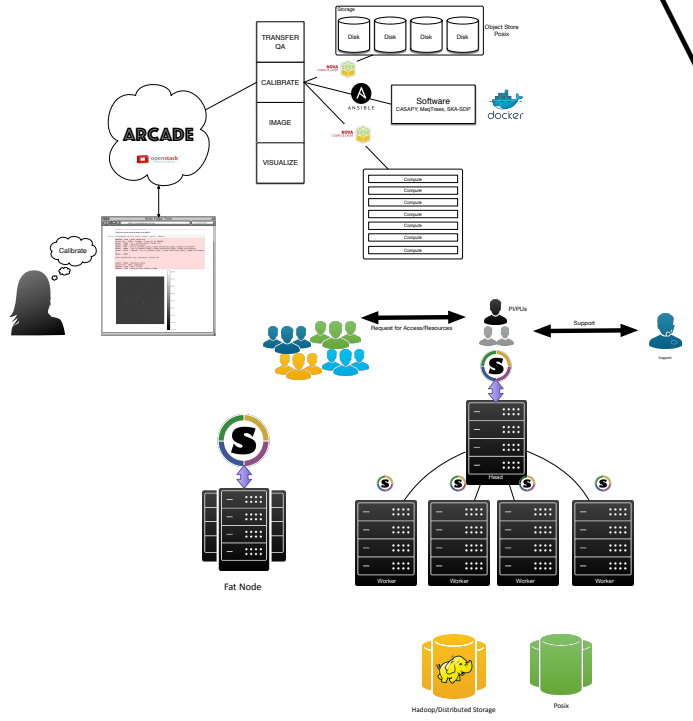
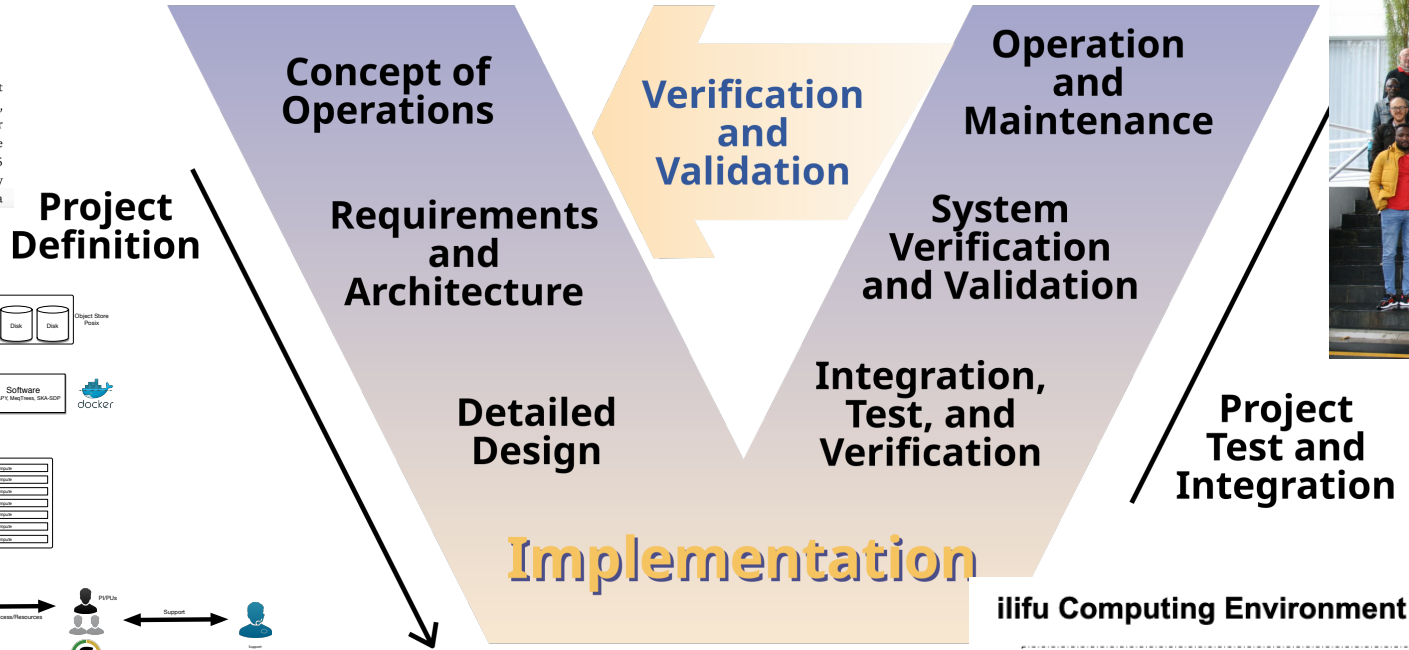
MeerKAT Science: On the Pathway to the SKA

MeerKAT2016

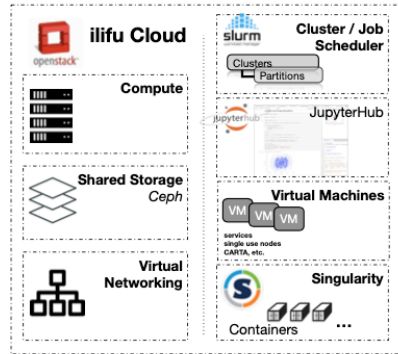
25-27 May, 2016
 Stellenbosch, South Africa
 published February 01, 2018

Entries on ADS

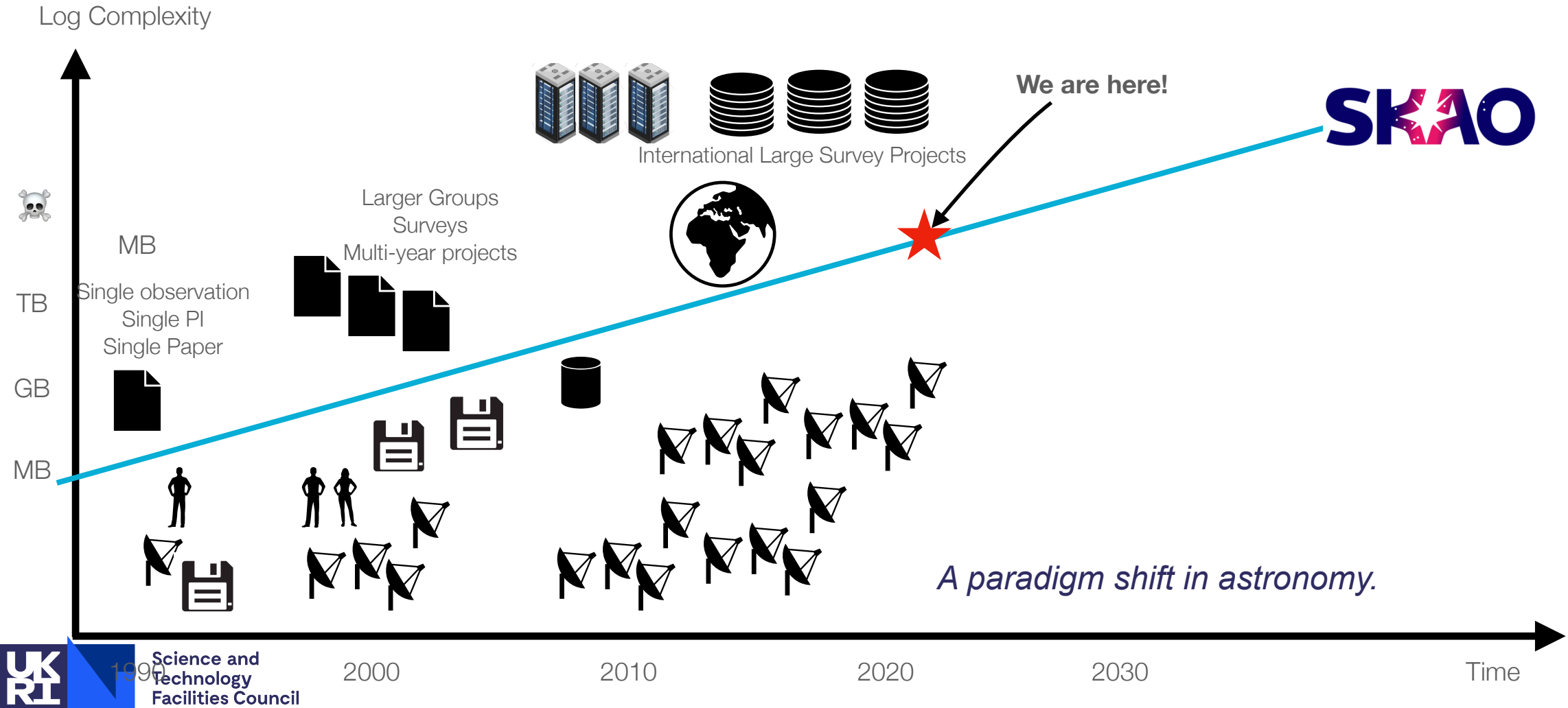
MeerKAT Science: On the Pathway to the SKA



ilifu Computing Environment



MeerKAT/IDIA in Context



Time to Science

Model of the data work flow

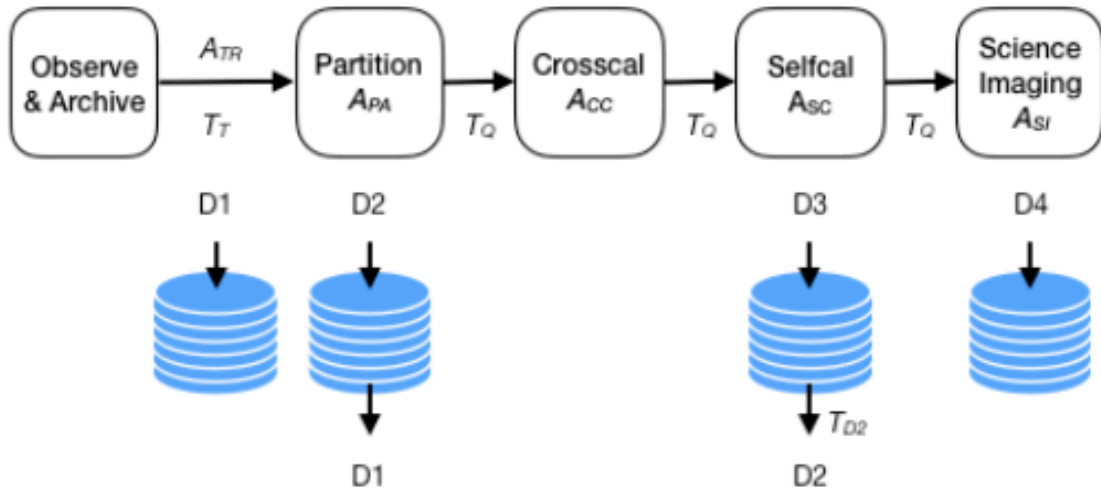
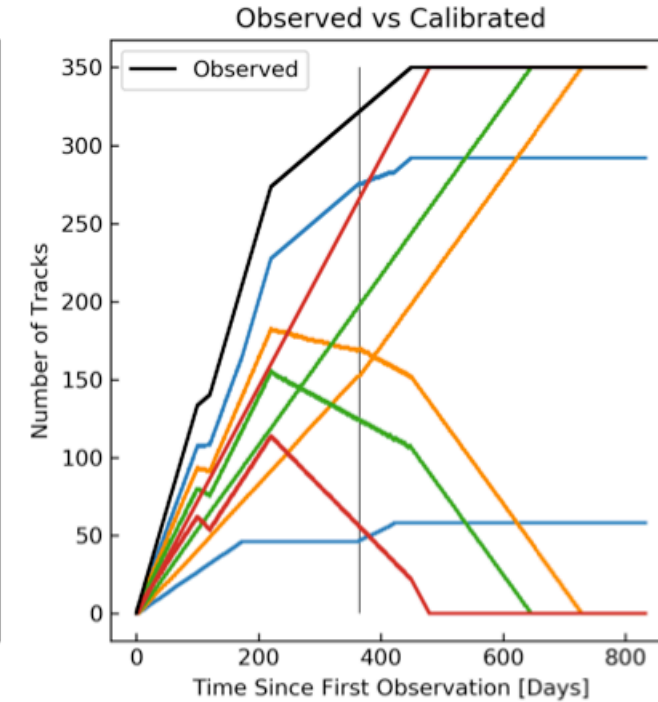
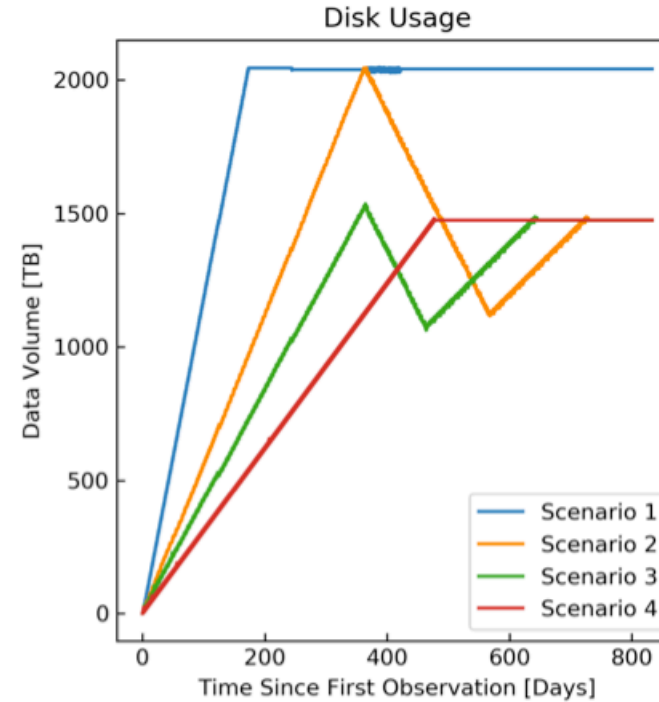


Figure 3: Overview of the LSP operational model at IDIA. Each block illustrates a processing step, and has an associated coefficient. The model includes the corresponding data write or removal operation, and various wait or staging times.

$$T_{proc} = (A_{TR} + A_{PA} + A_{CC} + A_{SC} + A_{SI}) \cdot T_{obs}$$

From Ilifu Processing and Storage
for MeerKAT Large Imaging Surveys (Memo)



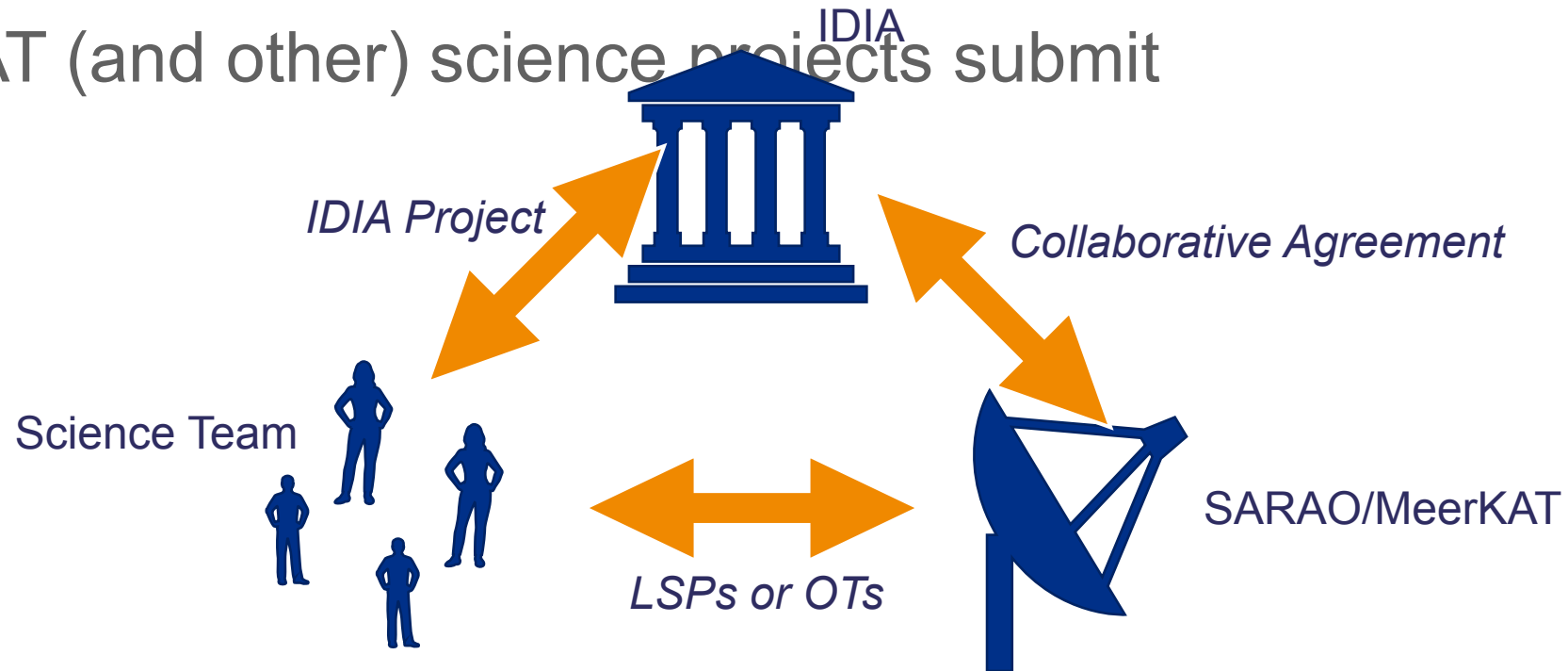
Each scenario explores the impact of:

- Serial vs parallelised processing
- Data production QA
- QA storage overhead (hoarding data)
- Science imaging strategies.

Implemented using SIMPY

“Governance”

- Independent, with an oversight committee comprising DVCs and SARA0 Chief Scientist.
- Funding — tapped from universities and national funding agency (NRF).
- MeerKAT (and other) science projects submit



IDIA and MeerKAT

- SARA0: call for proposals, or large project review, which includes a data processing/management plan or report.
- Simultaneously, IDIA issues a call for expressions of interest, or support, for data management.
- Letter of support, or *data/report*, is provided for potential projects, which is included in proposal, or report.
- For successful projects, IDIA+SARA0 sets-up semi-automatic data transfer for observations, and IDIA projects (fairshare rated).
- IDIA reviews projects annually, based on data usage.

Capability

Clear link between strategy (big picture) and project development.



ARC
 Proof of Concepts
 Small Teams
 Project Management

IDIA
 Procurement/Commissioning
 Developing
 Small Teams
 Project Management



IDIA/Ilifu/SARAO
 Operations
 Growth Focused
 Career Development
 People Management



- Operational needs are different!
- * Build and maintain high performance teams.
 - * Balance autonomy and hierarchy.
 - * Develop policies/best practices.
 - * Adhere to POPI/GDPR, Sorbonne Charter.
 - * Constant Training.
 - * Community engagement/training.

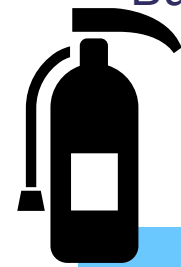
Time

Capability

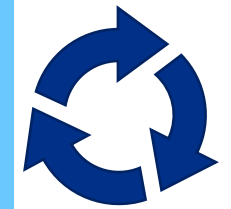
Clear link between strategy (big picture) and project development.



Business Development



IDIA/Ilifu/SARAO
Operations
Growth Focused
Career Development
People Management



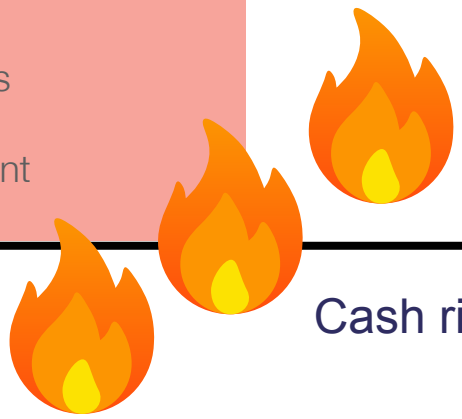
ARC
Proof of Concepts
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Operational needs are different!

- * Build and maintain high performance teams.
- * Balance autonomy and hierarchy.
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- * Adhere to POPI/GDPR, Sorbonne Charter.
- * Constant Training.
- * Community engagement/training.



Cash rich *burn* period

Time

Challenges

- Astronomers are not computers scientists!
- Policy development — flexible enough for innovative science, but strong enough to ensure compliance and fairness.
 - Also affects your technology, e.g., Slurm fairshare!
- Service at a *constant* level: $\epsilon_{\text{total}} = 2 \times \epsilon_{\text{OpEx}} \times T_{\text{ops}}$
 - Depends on Inflation, Operational Model, Staff, etc...
 - For IDIA, $\epsilon_{\text{OpEx}} \sim \text{€}0.5\text{M}$, so $\epsilon_{\text{total}} = \text{€}5\text{M}$ for 5 years
- Sustainable management model — scaling from a *project* to an institution.

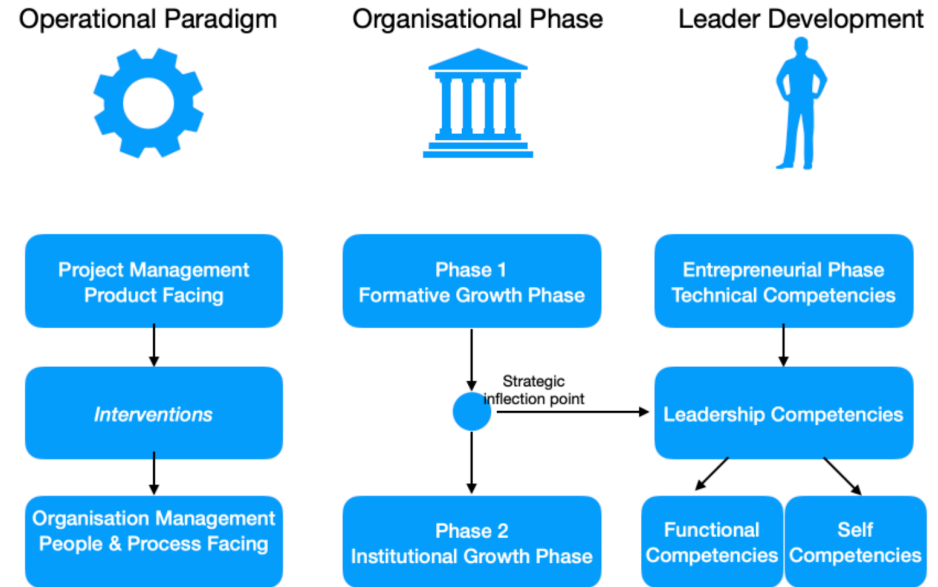


Table V

Self competencies

- A. Exercise intellectual integrity**
 1. Understand personal strengths and weaknesses
 2. Hire people to complement weaknesses
 3. Self educate where needed
- B. Move from “me” to “we”**
 1. Sell the company instead of self
 2. Become a coach
- C. Speak to the oracle**
 1. Maximize external advisors
 2. Participate in networking groups
 3. Anticipate organizational changes
- D. Create a sustainable organization**
 1. Become a strategic thinker
 2. Begin to make long-term decisions
 3. Identify unique contribution to the organization

